

# Statistical Study of Filament Eruptions and Moreton Waves Observed by the Flare Monitoring Telescope at Hida Observatory, Kyoto University

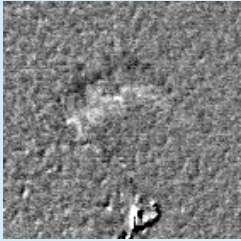


M. Yamaguchi, K. Shaltout, A. Asai, S. Morita, K. Shibata, R. Kitai, T. T. Shii, S. Ueno, N. Makamura, S. Takasao, Y. Yoshinaga, A. Hillier (Kyoto Univ.), K. Otsuji (NAOJ), N. Narukage (JAXA), D. Cabezas, M. Gutierrez, Y. Buleje, M. Ishitsuka, J. Ishitsuk (IGP), R. Terrazas, L. Martinez (Ica Univ.)

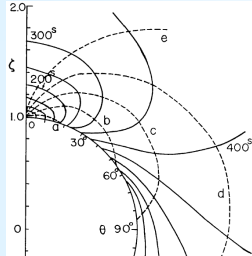
## INTRODUCTION

What is Moreton wave?

- Flare related wave-like phenomena in H $\alpha$
- V ~1000km/s, limited opening angle
- Interface between MHD fast mode shock (traveling in the corona) and the chromosphere (Uchida model)



Moreton wave with flare in November 4, 1997



Uchida Model

### Conditions of Moreton Wave

- Moreton waves are rare!
  - tend to be associated with large flares
  - always associated with filament eruptions (the direction of a Moreton waves is the same as that of the associated filament eruption)

[idea]

- Fast mode shock that generated Moreton waves are "piston-drive"?
- The shock front must be horizontal to intersect with the chromosphere (and to generate a Moreton wave) ?

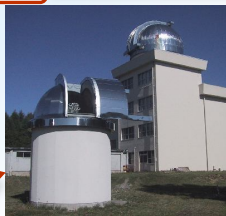
### motivation

To confirm the idea, we investigated elevation angles of filament eruptions statistically observed with Hida/FMT

## OBSERVATION

Instrument

- Flare Monitoring Telescope (FMT)
  - five refractors diameter 64mm
  - focal length 1920mm
  - wavelength
    - H $\alpha$  center +0.8A
    - 0.8A
    - continuum
    - H $\alpha$  for limb prominence



FMT

1992-2006, 2010 move to Peru

## DATA

Table 3.1: Event list of Moreton waves observed with the FMT.

| date       | peak time | flare position | NOAA AR | class | Moreton wave and/or winking filament(*) |
|------------|-----------|----------------|---------|-------|---|
| 1997/11/03 | 04:28     | S20 W13        | 8100    | C8.6  | MW                                      |
| 1997/11/04 | 06:02     | S14 W33        | 8100    | X2.1  | MW & WF                                 |
| 1998/08/08 | 03:17     | N17 E74        | 8299    | M3.0  | MW                                      |
| 1999/02/16 | 03:12     | S23 W14        | 8458    | M3.2  | MW & WF                                 |
| 2000/03/03 | 02:14     | S15 W60        | 8882    | M3.8  | MW                                      |
| 2000/06/04 | 22:10     | N21 E27        | 9026    | M3.2  | WF                                      |
| 2000/06/15 | 23:43     | N19 E19        | 9040    | M2.0  | MW                                      |
| 2000/07/16 | 06:14     | S08 W25        | 9082    | C3.8  | MW                                      |
| 2001/04/10 | 05:26     | S23 W09        | 9415    | X2.3  | WF                                      |
| 2001/05/12 | 23:35     | S17 E00        | 9455    | M3.0  | MW                                      |
| 2001/05/13 | 03:04     | S18 W01        | 9455    | M3.6  | MW                                      |
| 2001/05/13 | 23:09     | S15 W13        | 9455    | C1.1  | WF                                      |
| 2001/05/21 | 03:20     | N22 E08        | 9461    | C9.0  | WF                                      |
| 2001/12/19 | 02:32     | N09 E37        | 9742    | C4.9  | MW                                      |
| 2002/07/18 | 07:44     | N19 W30        | 0030    | X1.8  | WF                                      |
| 2002/08/22 | 01:57     | S07 W62        | 0069    | M5.4  | MW                                      |
| 2002/10/04 | 22:43     | N13 E43        | 139     | M2.7  | MW                                      |
| 2003/05/27 | 23:07     | S07 W17        | 365     | X1.3  | MW                                      |
| 2004/11/03 | 03:35     | N09 E45        | 0696    | M1.6  | MW                                      |
| 2005/08/03 | 05:06     | S13 E45        | 0794    | M3.4  | MW                                      |

We analysis 14 events from all 20 events  
(There are 374 X & M class flare in cycle 23(1996-2008))

## ANALYSIS

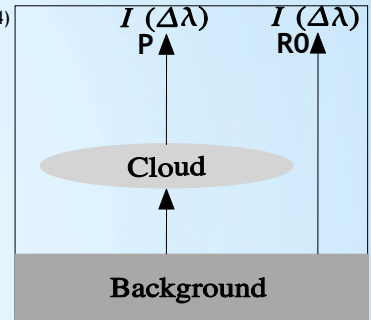
Analysis Method

- Beckers' Cloud Model (Beckers 1964)
- Contrast function C (x,  $\Delta\lambda$ )

$$C(x, \Delta\lambda) = \frac{I_P(x, \Delta\lambda) - I_{RO}(\Delta\lambda)}{I_{RO}(\Delta\lambda)}$$

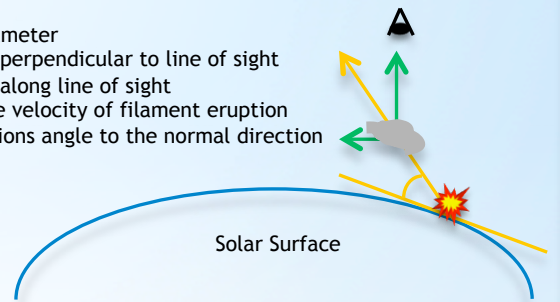
$I_p$ : Cloud intensity  
 $I_{RO}$ : Background intensity

From the contrast function, physical parameters of erupting filaments such as source function and doppler shift are determined



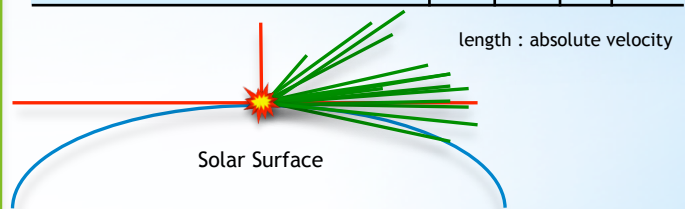
Obtained parameter

- the velocity perpendicular to line of sight
- the velocity along line of sight
  - absolute velocity of filament eruption
  - inclinations angle to the normal direction



## RESULT

| Date       | Peak time | Position | NOAAAR | Class | Moreton wave and/or winking filament(*) | Elevation angle | Absolute Speed | Doppler Velocity | Transversal Velocity |
|------------|-----------|----------|--------|-------|---|-----------------|----------------|------------------|----------------------|
| 1997/11/03 | 4:28      | S20 W13  | 8100   | C8.6  | MW                                      | -14             | 170            | 27               | 170                  |
| 1997/11/04 | 5:58      | S14 W33  | 8100   | X2.1  | MW & WF                                 | 11              | 200            | 19               | 200                  |
| 1998/8/8   | 3:17      | N17 E74  | 8299   | M3.0  | MW                                      | 9               | 50             | 26               | 50                   |
| 1999/2/16  | 3:12      | S23 W14  | 8458   | M3.2  | MW & WF                                 | 45              | 100            | 28               | 100                  |
| 2000/3/3   | 2:14      | S15 W60  | 8882   | M3.8  | MW                                      | 43              | 140            | 26               | 130                  |
| 2000/6/15  | 23:43     | N19 E19  | 9040   | M2.0  | MW                                      | -6              | 160            | 25               | 150                  |
| 2000/7/16  | 6:14      | S08 W25  | 9082   | C3.8  | MW                                      | 30              | 110            | 8                | 110                  |
| 2001/5/12  | 23:35     | S17 E00  | 9455   | M3.0  | MW                                      | 20              | 160            | 6                | 160                  |
| 2001/5/13  | 3:04      | S18 W01  | 9455   | M3.6  | MW                                      | 53              | 30             | 17               | 20                   |
| 2001/12/19 | 2:32      | N09 E37  | 9742   | C4.9  | MW                                      | 8               | 170            | 26               | 170                  |
| 2002/8/22  | 1:57      | S07 W62  | 69     | M5.4  | MW                                      | 11              | 150            | 24               | 150                  |
| 2002/10/4  | 22:43     | N13 E43  | 139    | M2.7  | MW                                      | -37             | 170            | 7                | 170                  |
| 2003/5/27  | 23:07     | S07 W17  | 365    | X1.3  | MW                                      | 25              | 60             | 19               | 60                   |
| 2005/8/3   | 5:06      | S13 E45  | 794    | M3.4  | MW                                      | 4               | 130            | 22               | 130                  |



## DISCUSSION and SUMMARY

- The average value of absolute velocity summary ~130km/s(30-200km/s)
- The direction of eruption is almost horizontal (0~30°)
- This is follow the idea horizontal filament eruption drives Moreton wave
- We researched filament eruptions associated with Moreton wave and as the result, filament eruptions tend to be horizontal
- Now, We are studying the filament eruptions not associated with Moreton wave to compare