

0

-0.002

-0.004

-0.006

0

we define "Sun-like stars" (5600 K  $< \rm T_{eff} < 6000$  K, log g > 4.0, P<sub>rot</sub> >10 day). 44 superflares are observed on 19 Sun-like stars. The superflare frequency distribution on Sun-like stars is roughly on the same slope as that of solar flare, and locates between that of solar maximum and minimum. Stars more similar to the Sun (P<sub>rot</sub> >20 day) have superflare (Table below)



Superflares on Sun-like Stars with  $P_{rot} > 20$  days

KeplerID	$T_{\rm eff}$ (K)	$\logg$	$R/R_{\odot}$	$P_{\rm rot}(day)$	# of flares
5522535	5732	4.3	1.3	20.3	1
7597685	5834	4.6	0.9	21.8	1
8212826	5811	4.2	1.4	26.3	2
9766237	5674	4.6	0.9	21.8	1
9944137	5725	4.6	0.8	25.3	1
11401109	5732	4.5	0.9	29.1	1
Sun	5780	4.4	1.0	28.0	

Number of Flares, Flare Stars and Sample, and Fraction of Flare Stars

Slowly Rotating $(P_{\rm rot}>10d)$			Rapid	Rapidly Rotating $(P_{\rm rot} < 10d)$			Total					
$T_{\rm eff}$	$N_{\mathrm{flare}}$	$N_{\mathrm{flarestar}}$	$N_{\mathrm{all}}$	$f_{ m flarestar}$	$N_{\mathrm{flare}}$	$N_{\mathrm{flarestar}}$	$N_{ m all}$	$f_{ m flarestar}$	$N_{\mathrm{flarestar}}$	$N_{\rm all}$	$f_{ m flarestar}$	
5100-5600	353	50	14026	0.0036	810	133	1281	0.104	1163	183	15307	0.012
5600-6000	44	19	14325	0.0013	340	77	1825	0.042	384	96	16150	0.0059
	397	69	28351	0.0024	1150	210	3106	0.068	1547	279	31457	0.0089

Light curve of the Sun in 2003 (SORCE). Darkening caused by sunspot, which is similar to stellar light curve. This implies existence of stellar spot on superflare stars. Most of superflare stars show spotlike brightness variation. The typical amplitude is 1%, therefore, the typical spot size of superflare stars is 10% of stellar radius. We estimated stellar rotation period from brightness variation period.

0.3

TSI to the

20

Day

30

40

10

Days from flare peak Enlarged light curve of a superflare on the superflare star above. The amplitude of this superflare is about 0.3 % of stellar brightness. Typical amplitude of superflare is 1 % of stellar brightness. This is larger than that of solar flare detected from total solar irradiance (0.03%)

0.2

0

## Superflare on the Sun?

Our study suggests a occurrence of superflare on the Sun. <u>Typical frequency</u> of superflare of 10<sup>34</sup> erg on Sun-like stars is once in 2000 years. From theoretical point of view, our Sun can store enough magnetic energy to cause superflare of 10<sup>34</sup> erg within one solar cycle (Shibata et al. 2013). Observationally, large cosmic ray events in 7th and 9th century are found from tree ring (Miyake et al.). Although the source of this cosmic ray is under discussion, this event can be caused by solar superflare. The frequency of these events is consistent with superflare frequency of this study. In 1989, X15 solar flare caused large scale black out in Quebec, Canada and broke a transformer in USA. If solar superflare occur in this era of electronic communication, estimated damage is very large. Black out and communication failure all over the world and break down of almost all artificial satellite are possible.

## Reference

Shibayama, T., et al. 2013, ApJS, 209, 5 Schaefer, B. E., et al. 2000, ApJ, 529, 1026 Shibata, K., et al. 2013, PASJ, 65, 49 Mivake, F., et al. 2012, Nature, 486, 240

Maehara, H., et al. 2012, Nature, 485, 478 Notsu, Y., et al. 2013, ApJ, 771, 127 Notsu, S., et al. 2013, PASJ, 65, 112 Mivake, F., et al. 2013, NatureCom, 4, 1748



Takuya Shibayama<sup>1</sup> (shibayama@kwasan.kyoto-u.ac.jp),

Hirovuki Maehara<sup>2</sup>. Shota Notsu<sup>1</sup>. Yuta Notsu<sup>1</sup>. Satoshi Honda<sup>3</sup>. Daisaku Nogami<sup>1</sup>. Kazunari Shibata<sup>1</sup>

<sup>1</sup> Kyoto University, Kwasan Observatory; <sup>2</sup> University of Tokyo, Kiso Observatory; <sup>3</sup> Hyogo University, Nishi-Harima Observatory

Space weather study becomes increasingly important for our electronic civilization. One of the main cause of space environmental disturbance is large solar flare. The largest solar flare that our civilization have experienced is Carrington event in 1859, and the total energy of the flare is of the order of  $10^{32}$  erg.

We study superflares (whose total energy is more than  $10^{33}$  erg) on solar type stars with

0.002

-0.002

-0.004

0

0 2 4 6 8 10

Normalized



2012年(平成24年)5月28日(月曜日)

作のソフト