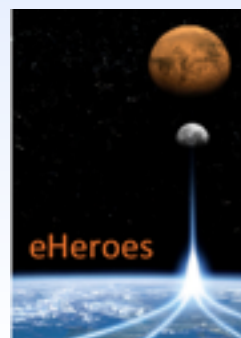


The topology of supersonic up-flows in a filament eruption

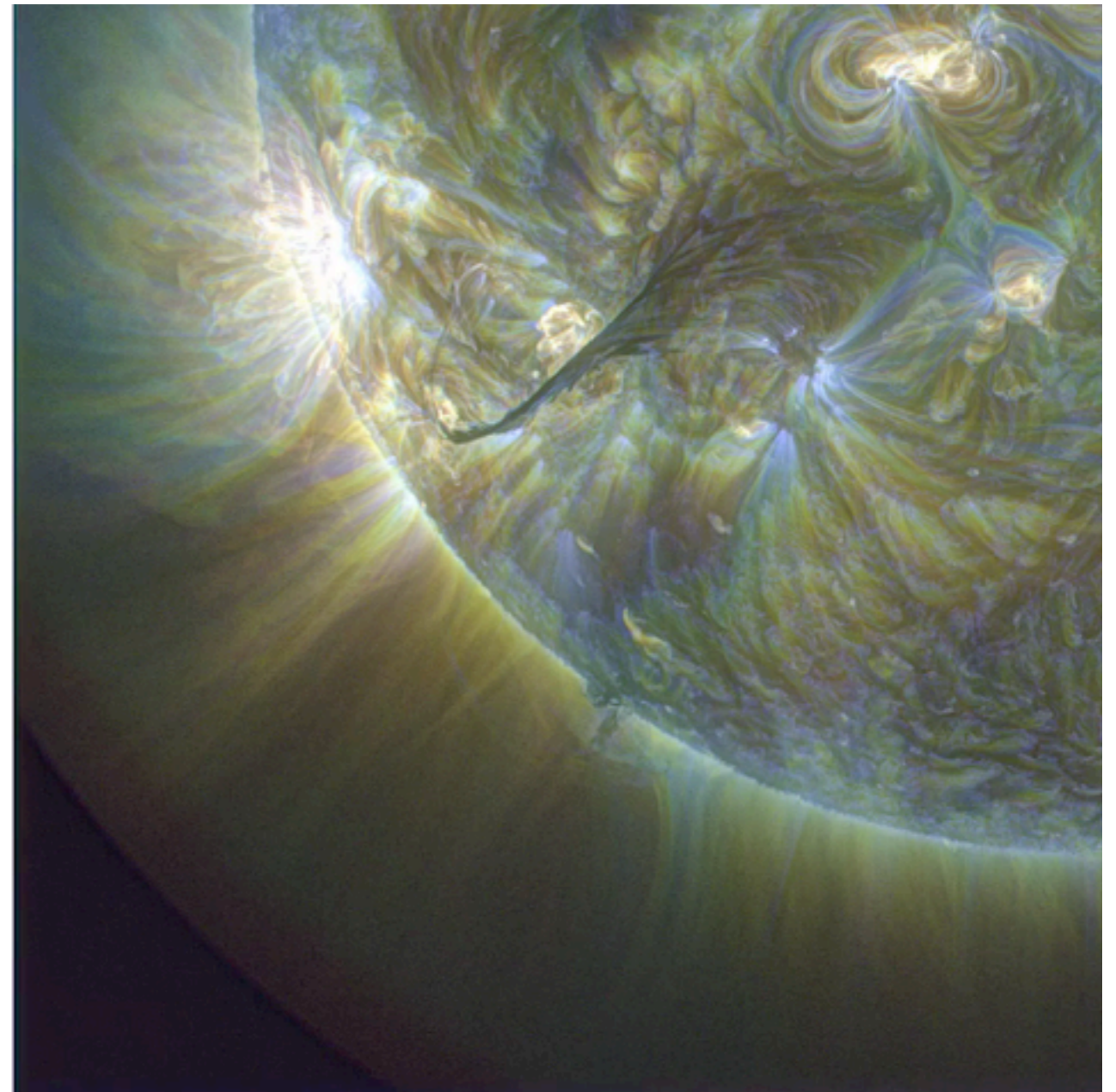
High-speed Flows in a Bright Thread

David Williams¹ • Deb Baker¹ • Lucie Green^{1,2} • David Long¹
 Lidia van Driel-Gesztelyi^{1,3,4} • Iain Hannah^{2,5}

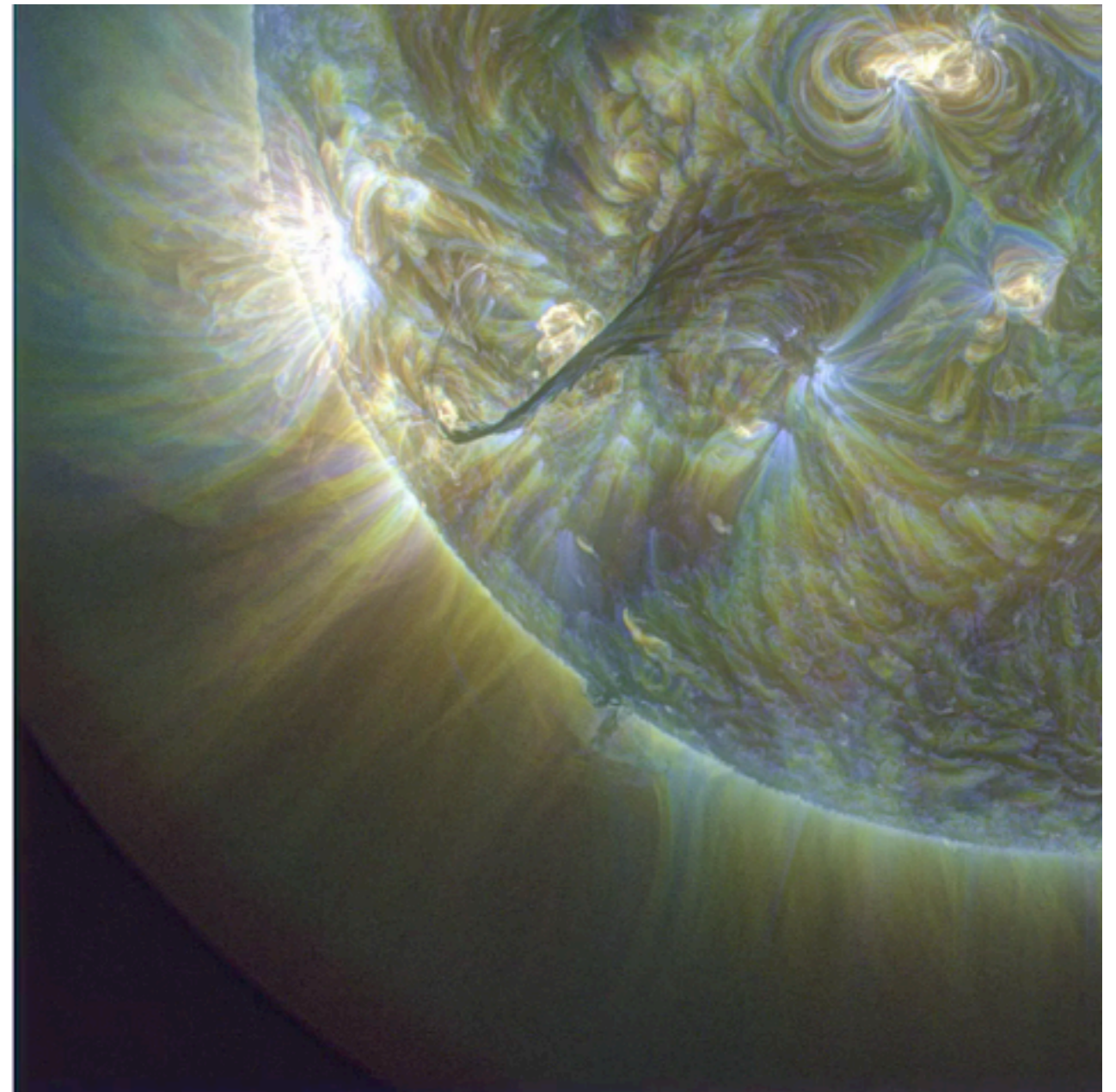
¹ University College London • ² Royal Society University Research Fellow
³ Observatoire de Paris, Meudon • ⁴ Konkoly Observatory
⁵ University of Glasgow



Filament eruptions & EIS

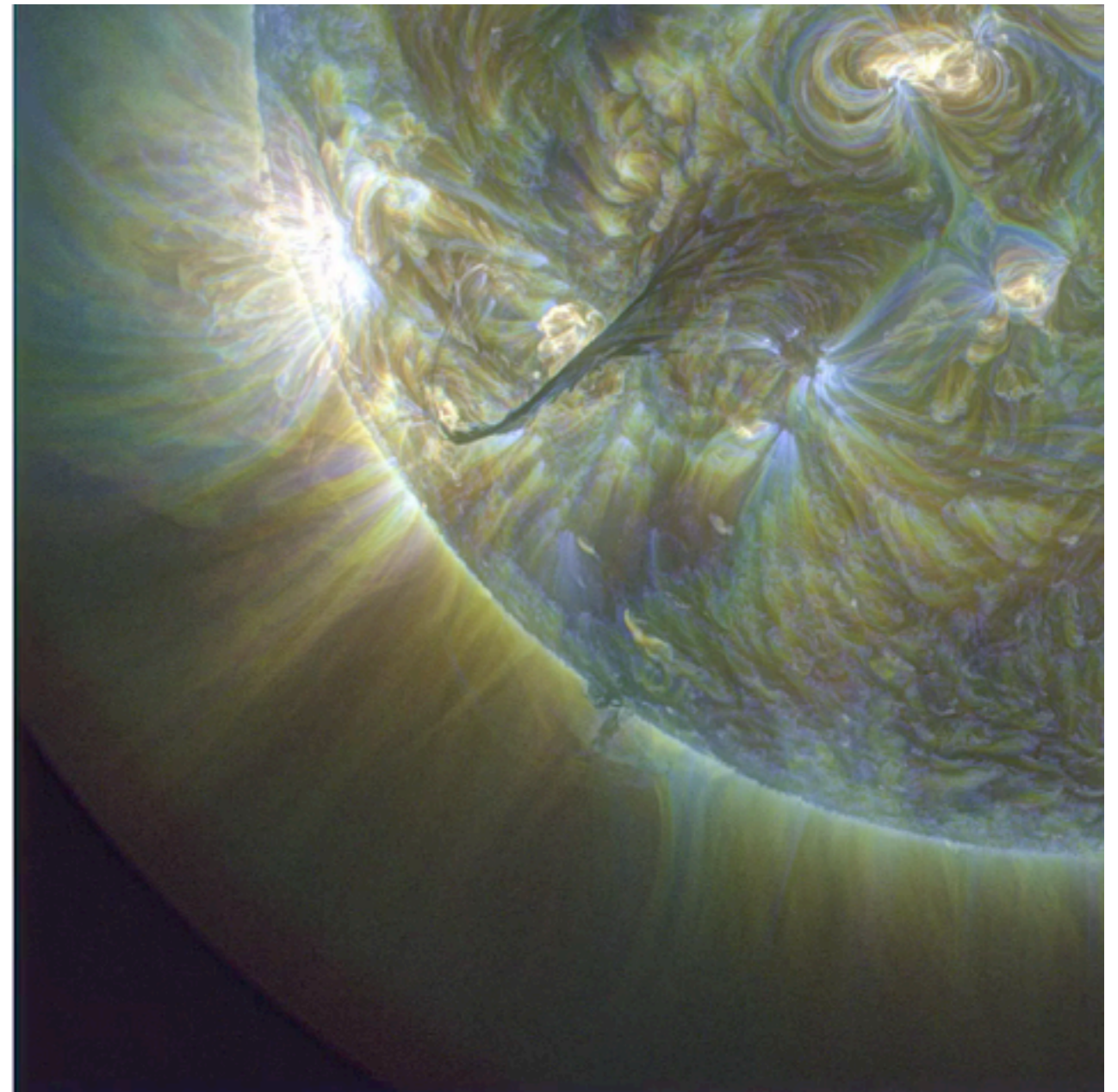


Filament eruptions & EIS



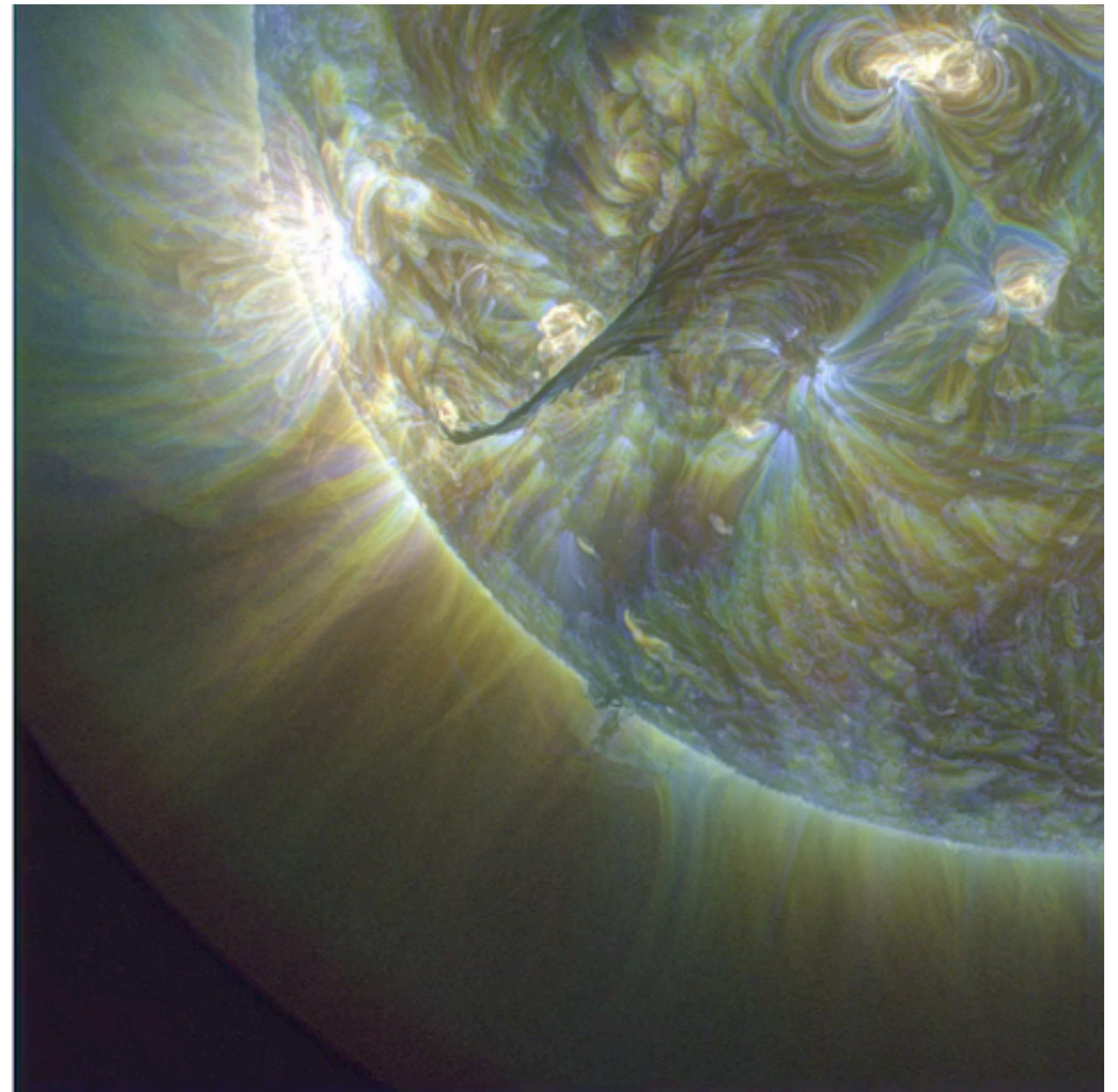
Filament eruptions & EIS

- Observed by AIA
 - **morphology**



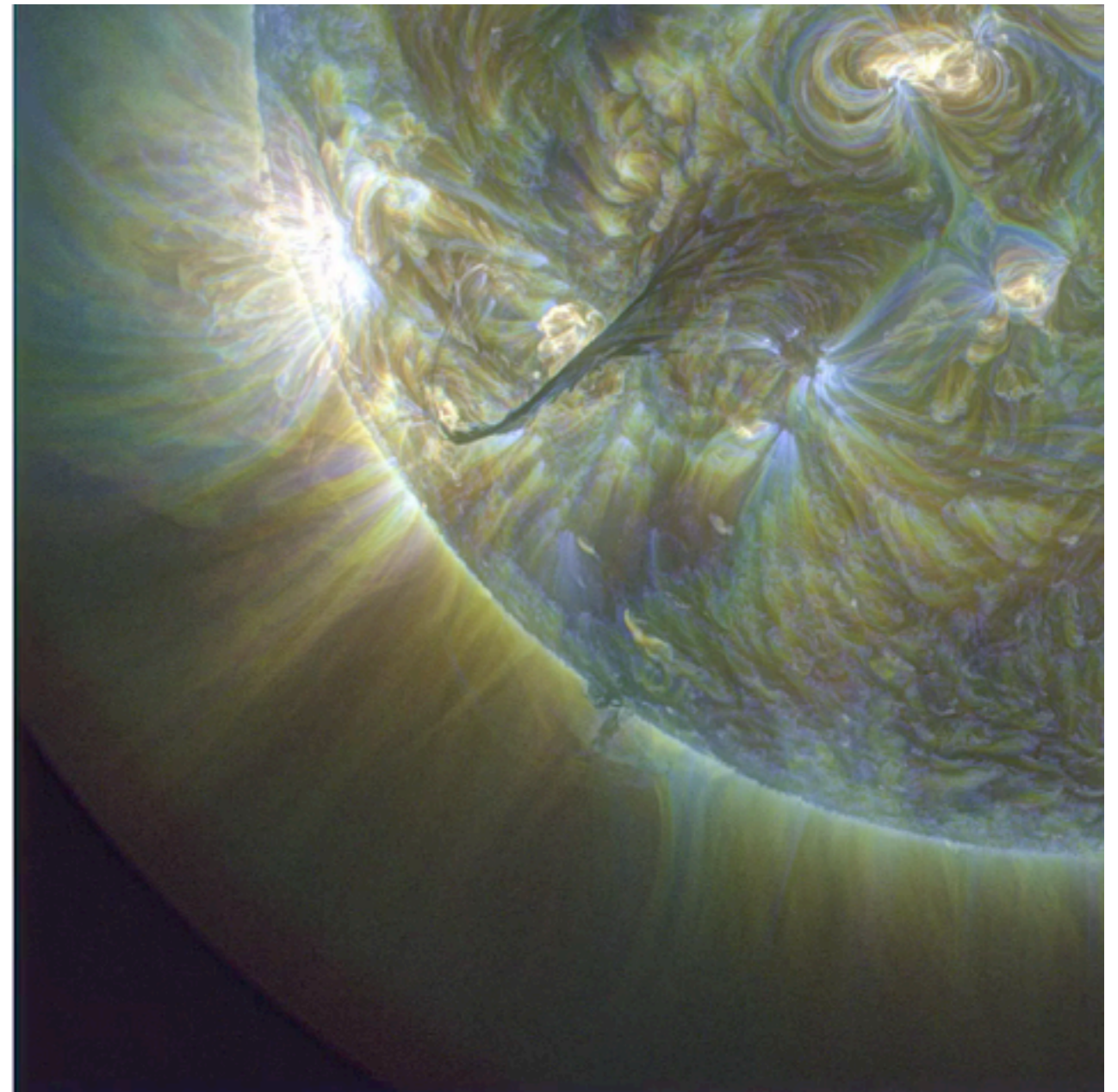
Filament eruptions & EIS

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 - other diagnostics



Filament eruptions & EIS

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 - **morphology**
- Observed by EIS
 - **velocity**
 - other diagnostics
- Probing dynamics
 - where eruption crosses EIS slit



Filament eruptions & EIS

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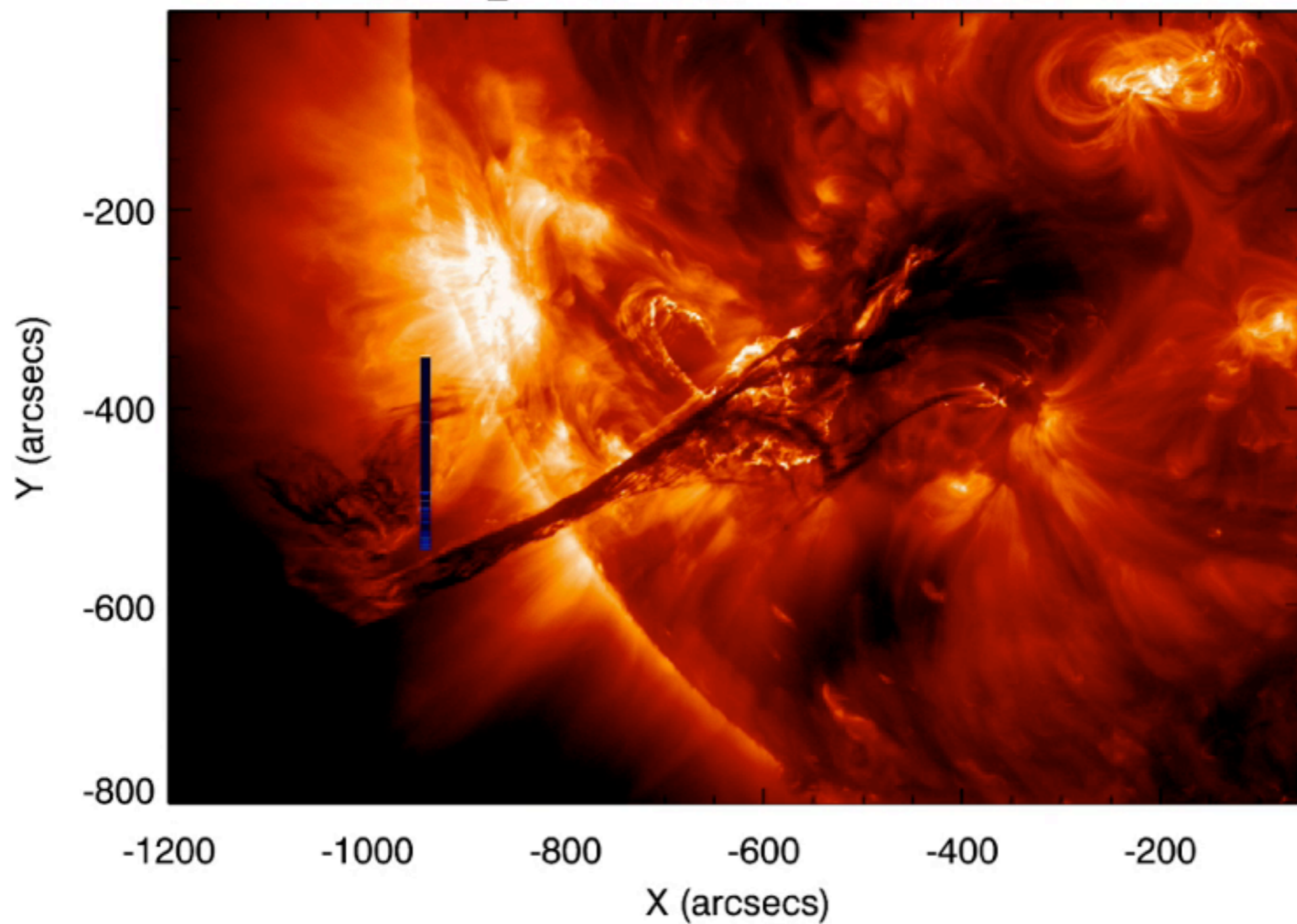
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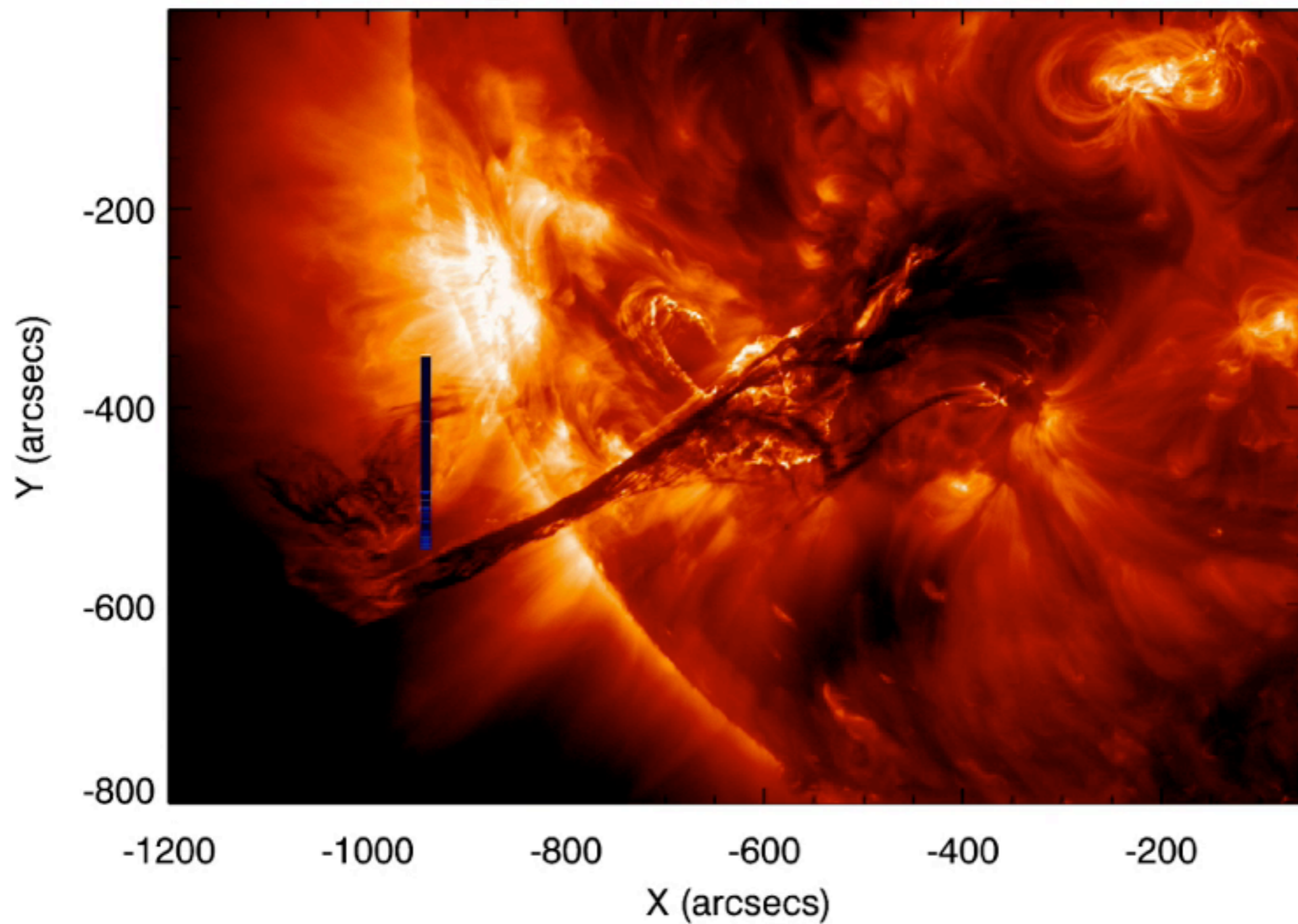
Filament eruptions & EIS

- Rare to be captured by satellite spectrometers
 - Small instantaneous field-of-view ($\leq 2'' \times 512''$)
 - Difficult to predict time and location
 - Spacecraft observing schedules of order days
 - Eruption pre-cursors on much shorter timescales

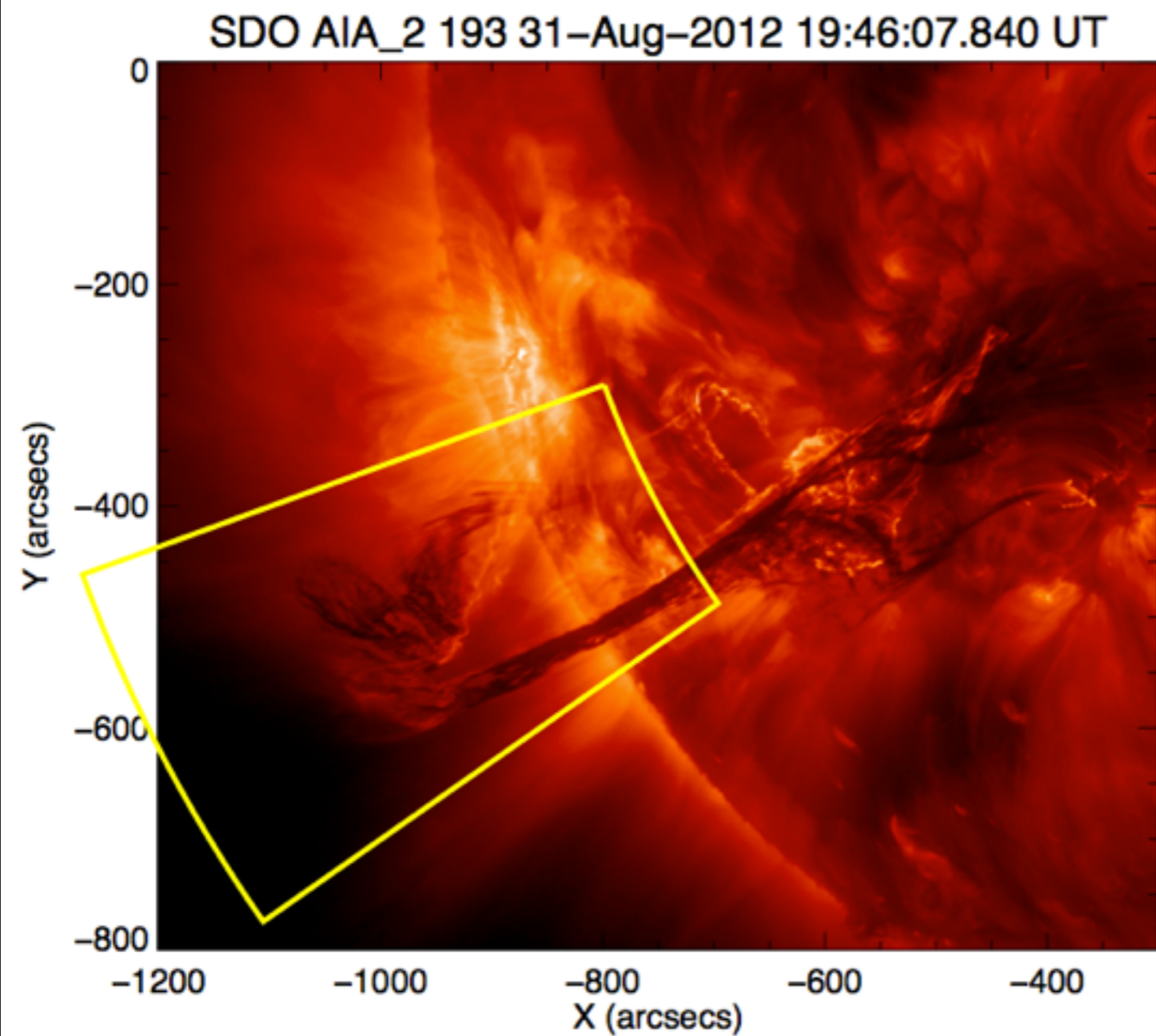
SDO AIA_2 193 31-Aug-2012 19:47:07.840 UT



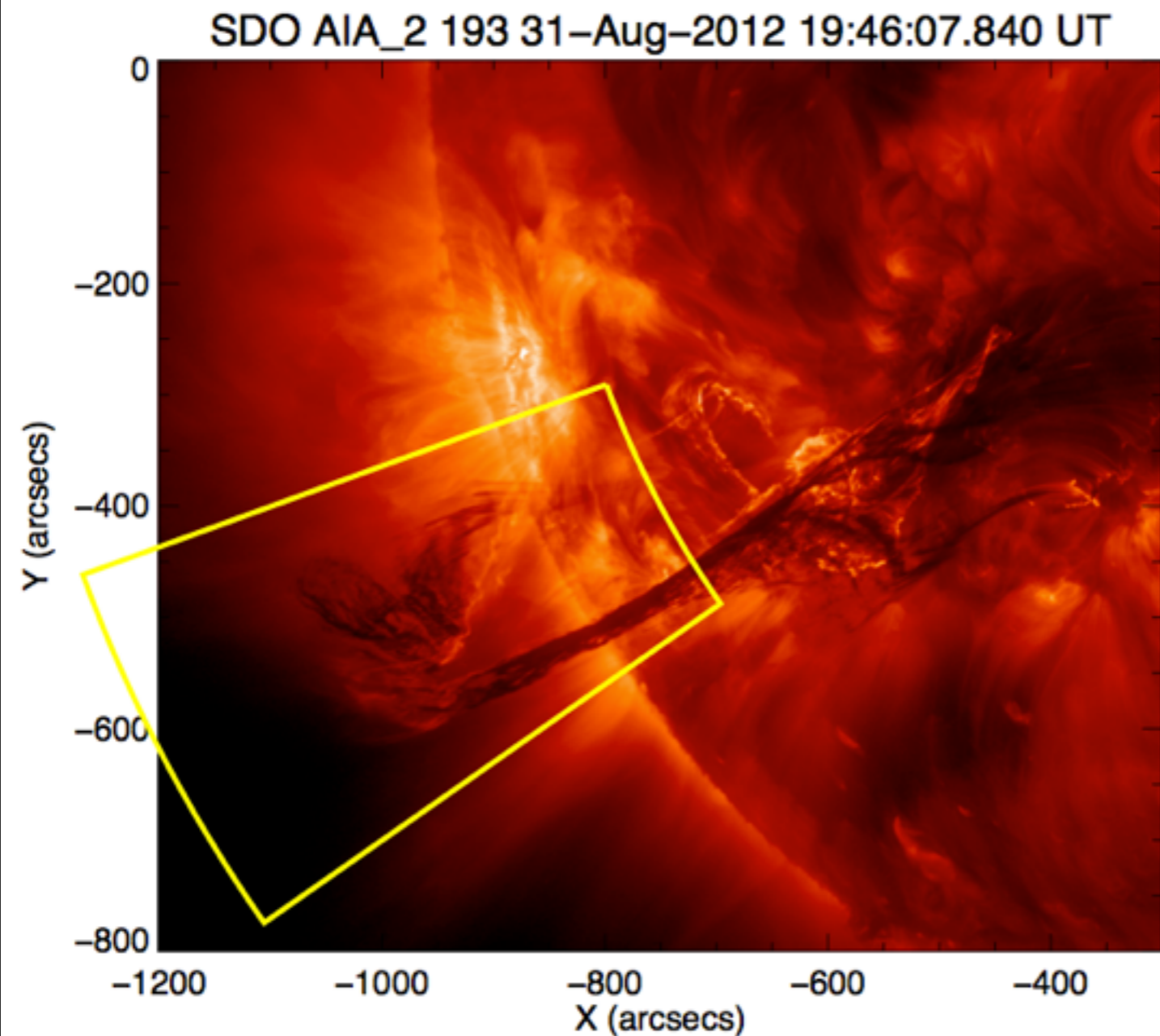
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Eruption profile near Sun

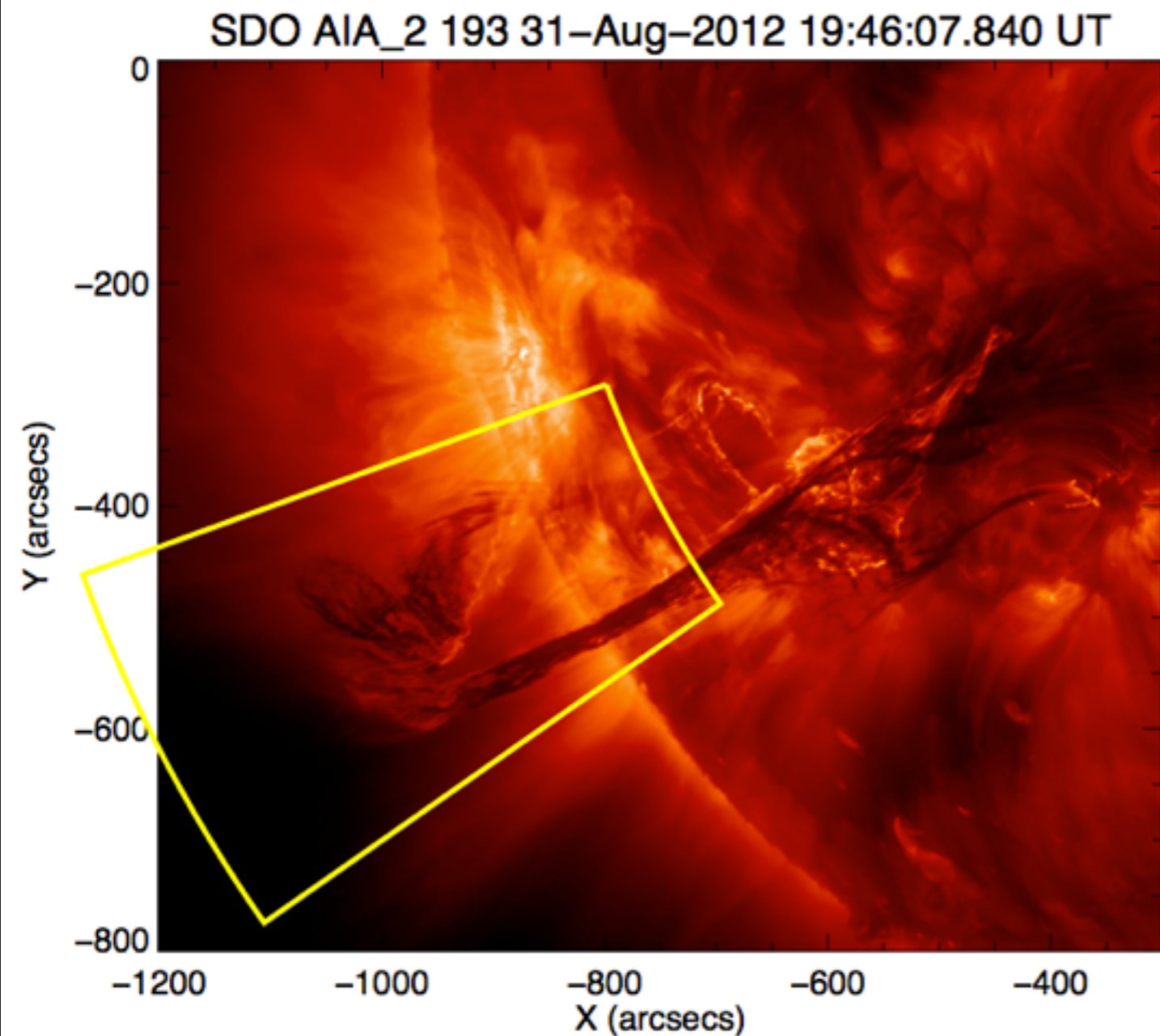


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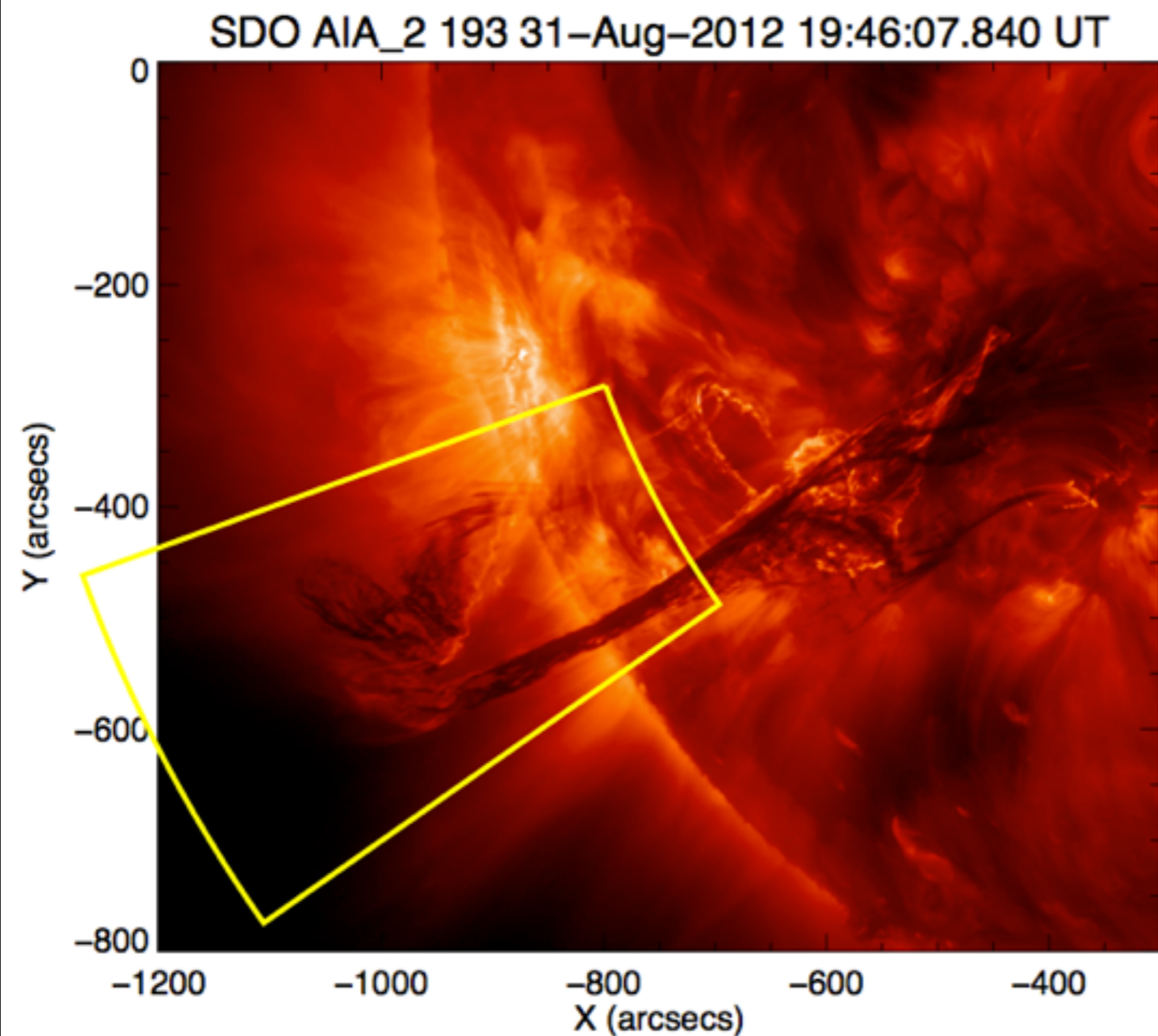
- Track radial motion of filament

Eruption profile near Sun



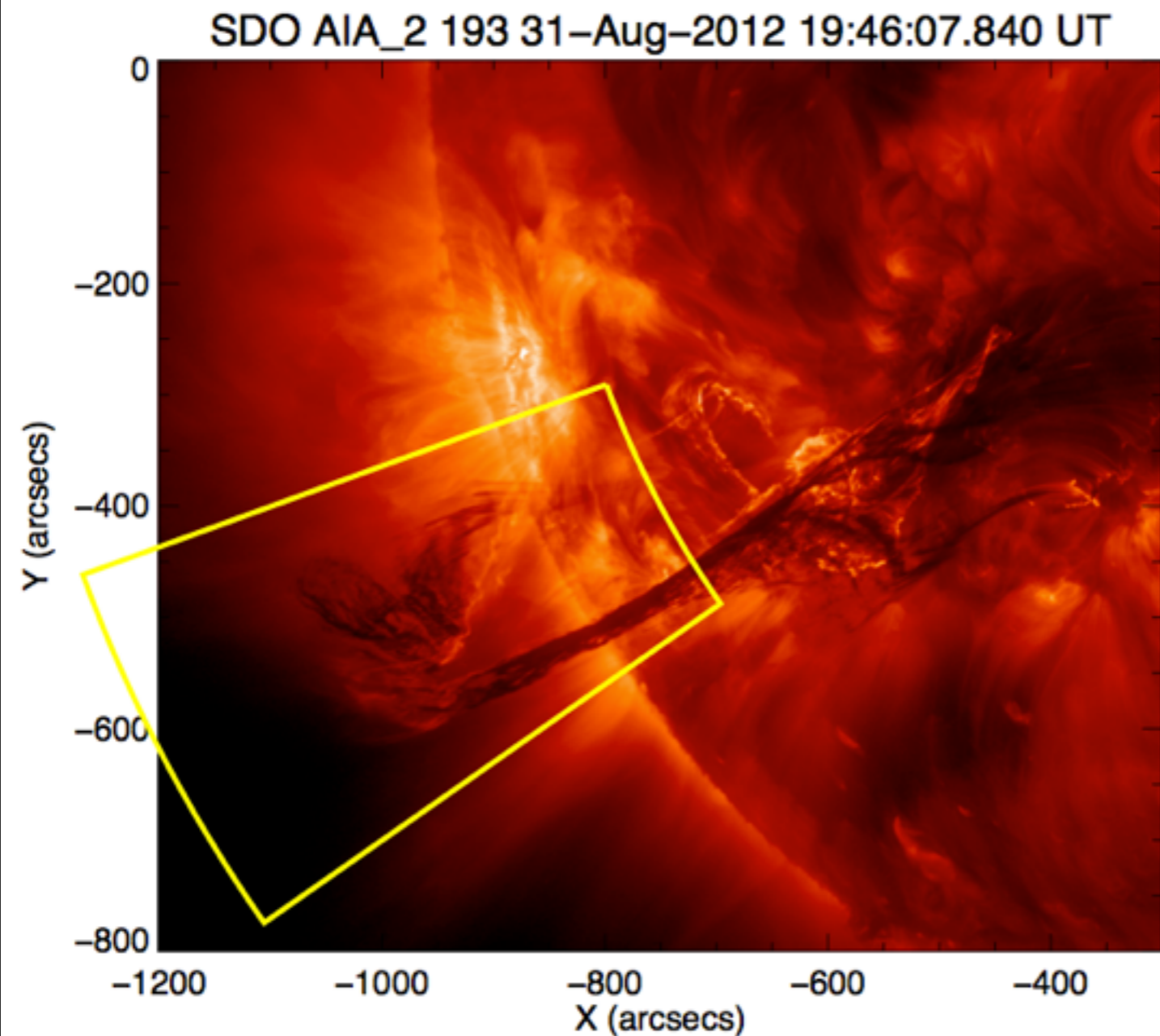
- Track radial motion of filament
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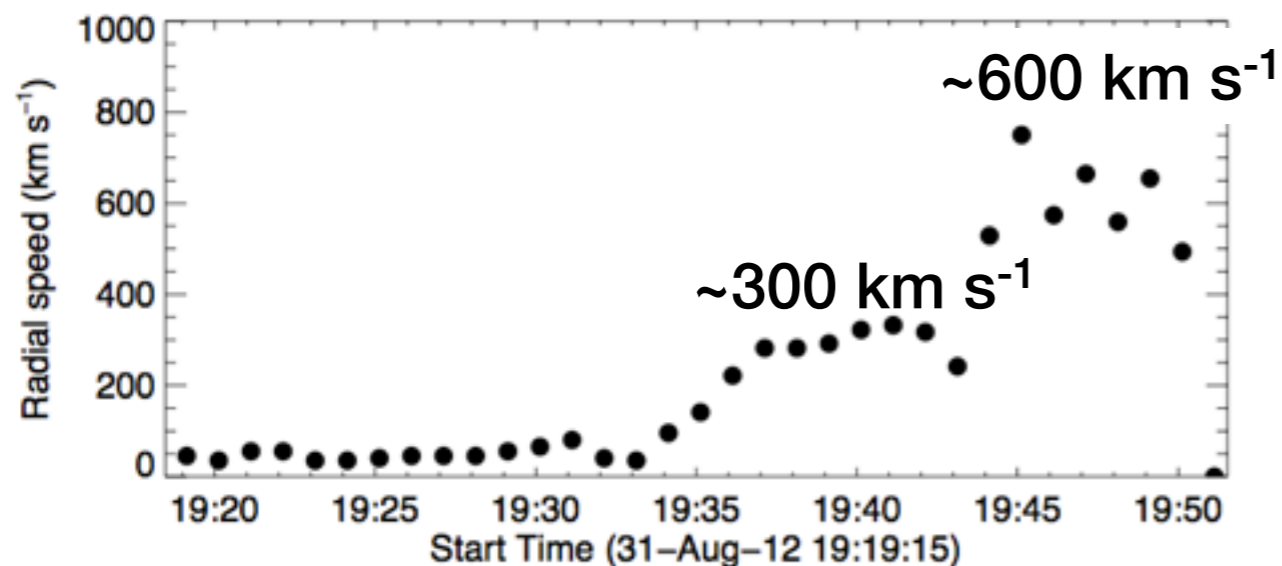
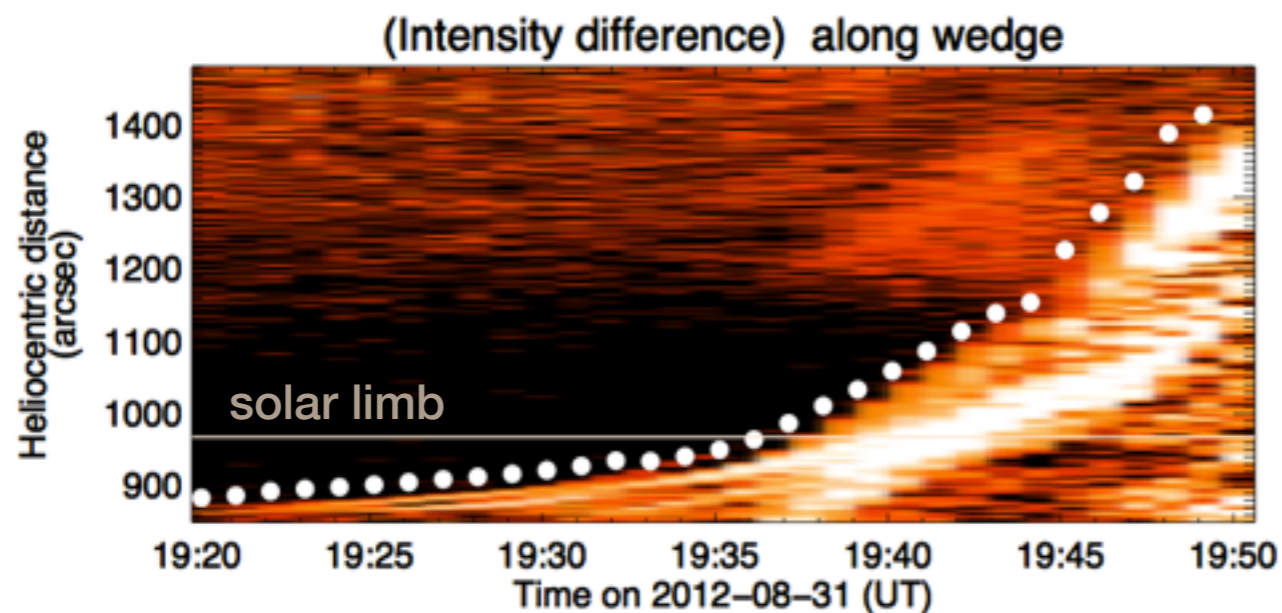
- Track radial motion of filament
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Eruption profile near Sun



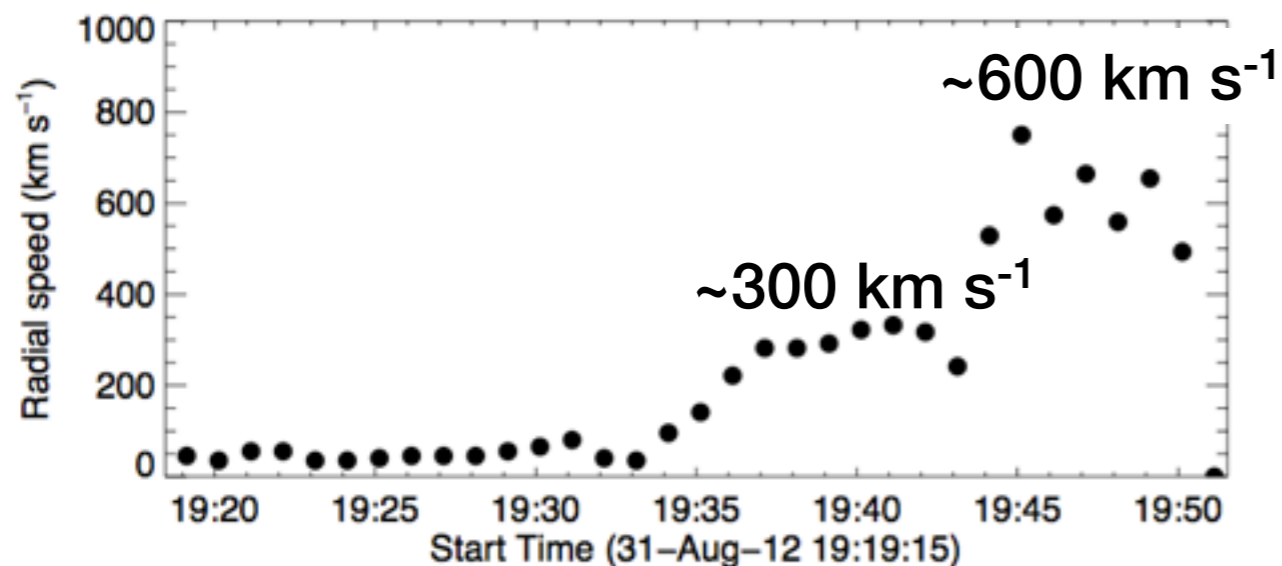
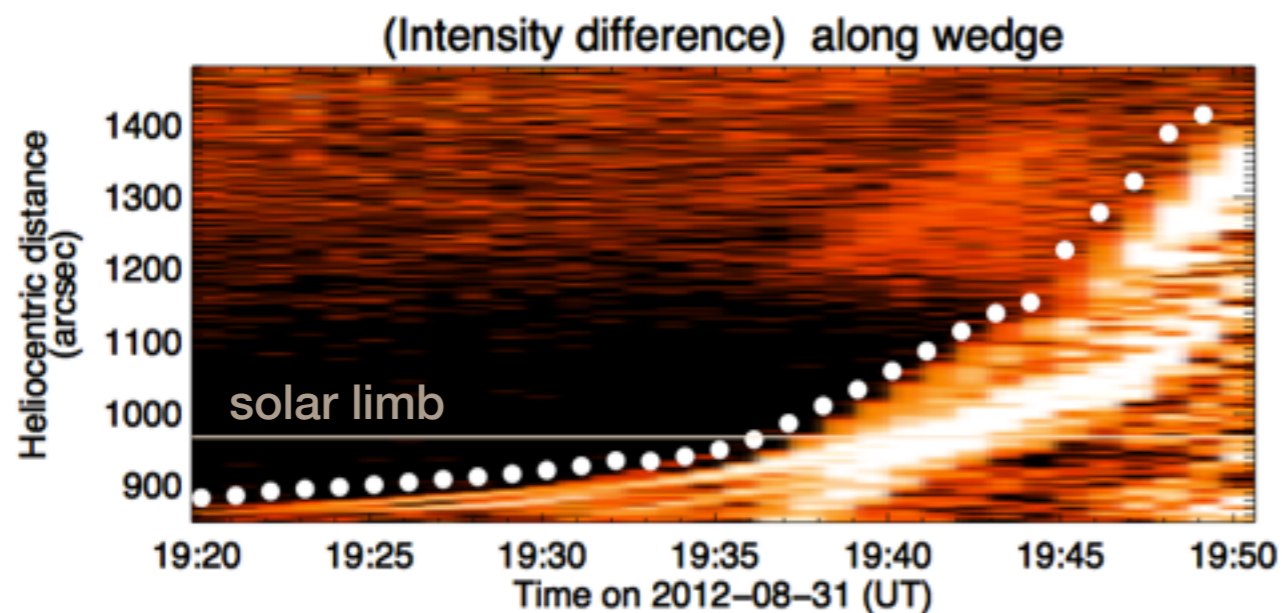
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Eruption profile near Sun



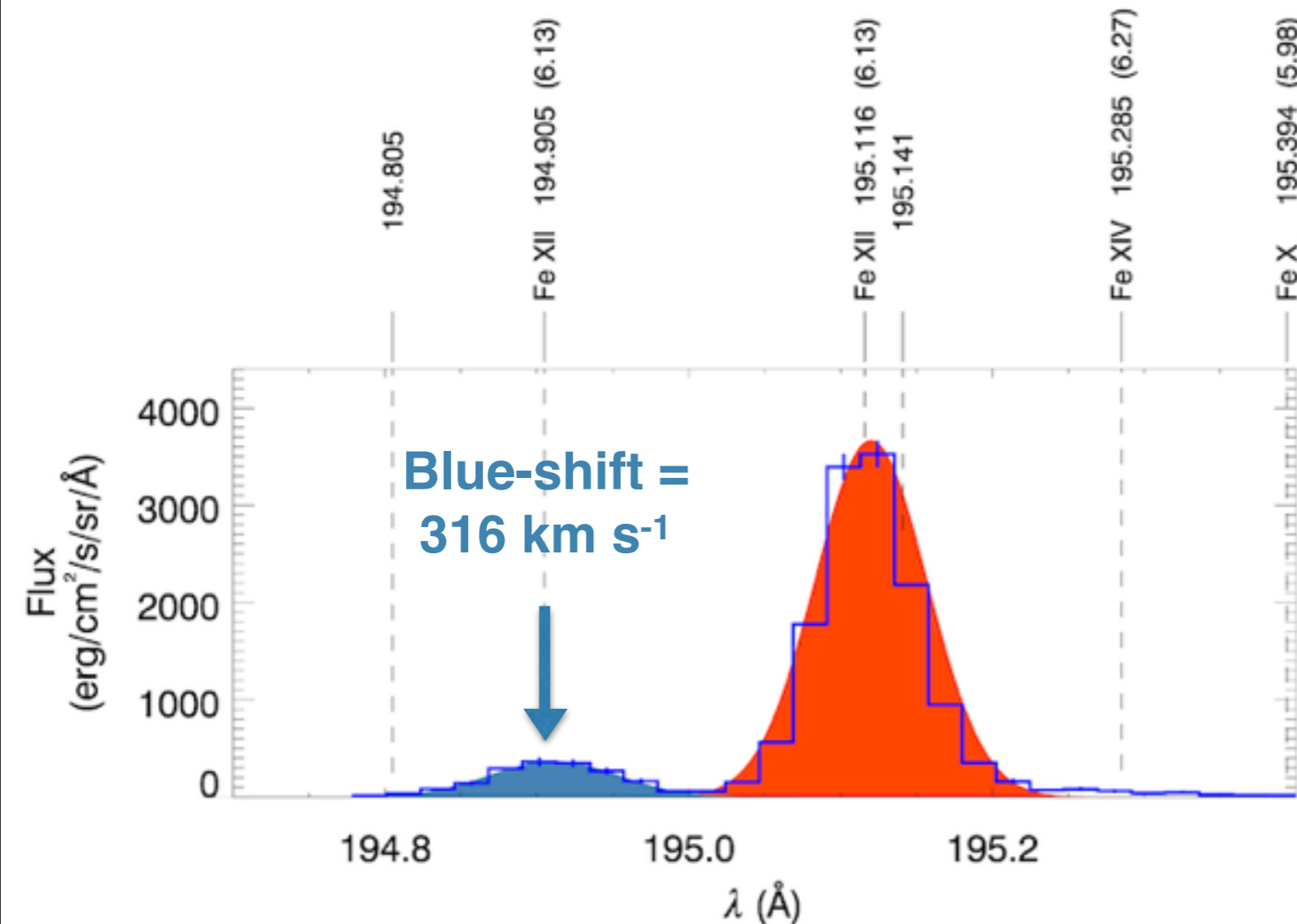
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Eruption profile near Sun



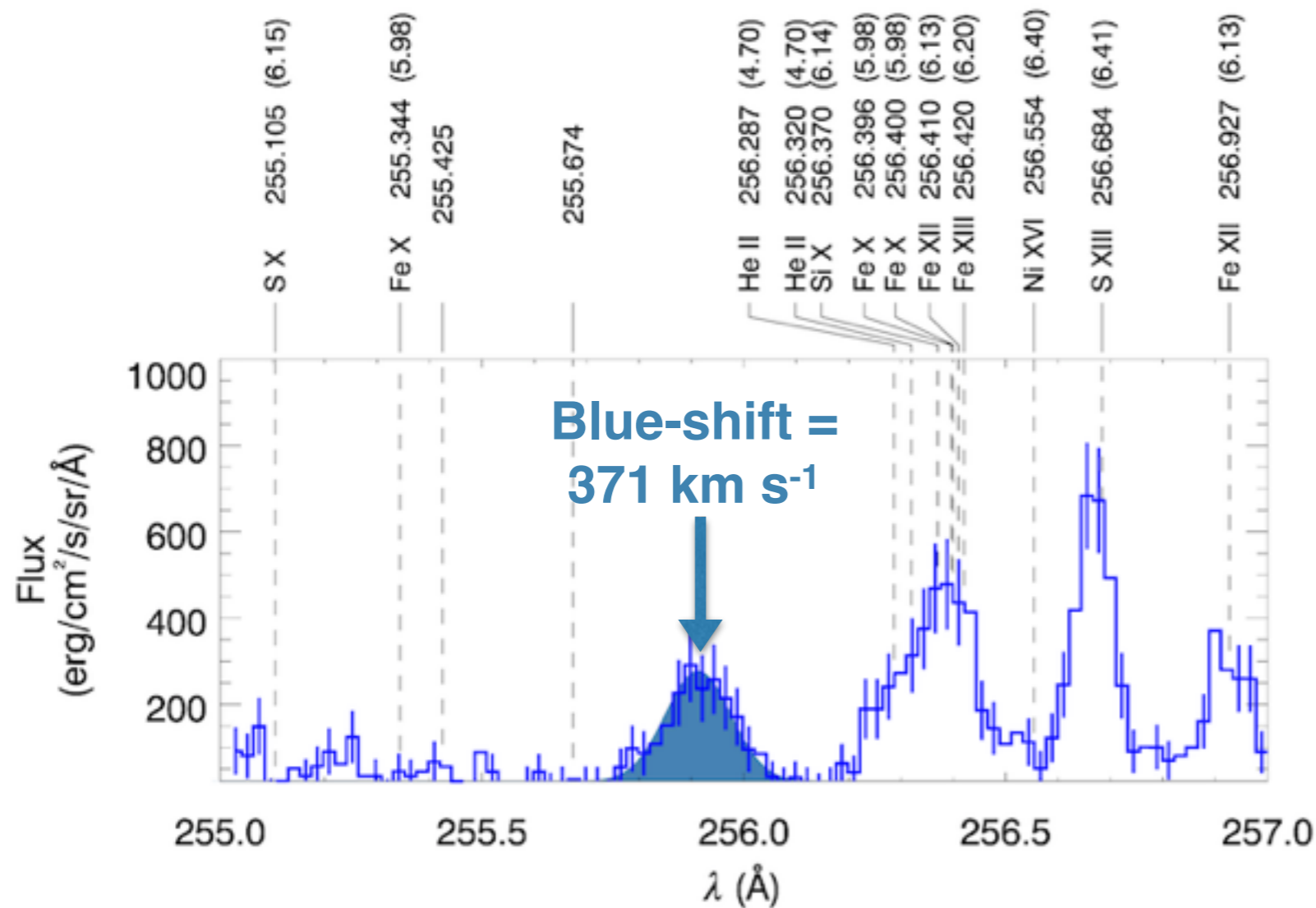
- Track radial motion of filament
- Filament shows EUV absorption and emission
 - Sum absolute differences from background
 - Integrate over 15° azimuth
 - Gradient-based tracking of front
- Shown as white circles

Spectra along filament: Fe XII



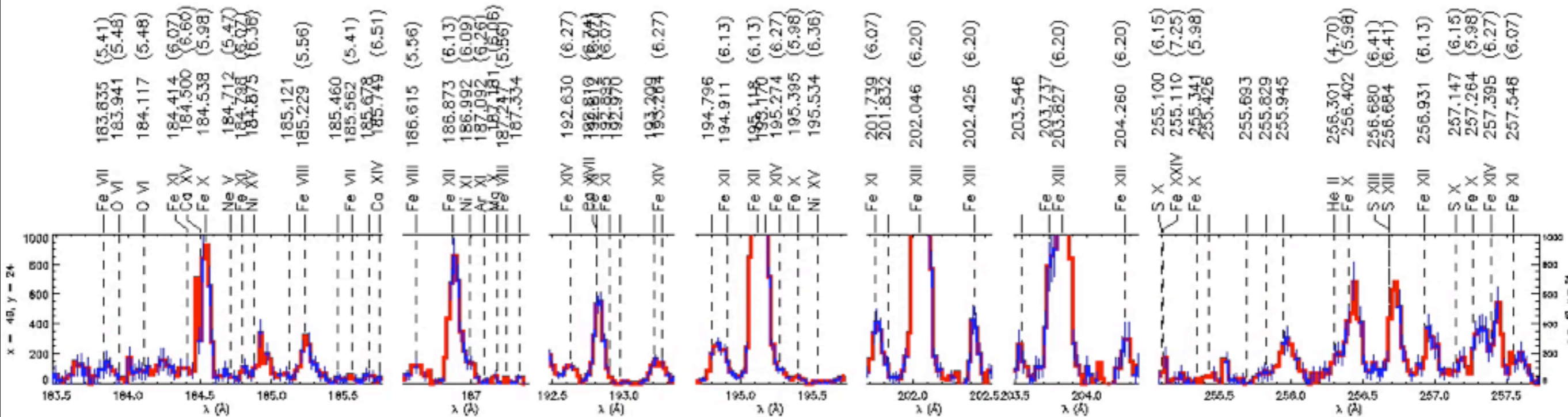
- 195.12 Å
- Rest component from foreground/background is shown in orange.
- No other strong lines in the spectrum where it shifts to: clean identification
 - λ194.9 is a red herring
- Supersonic

Spectra along filament: He II



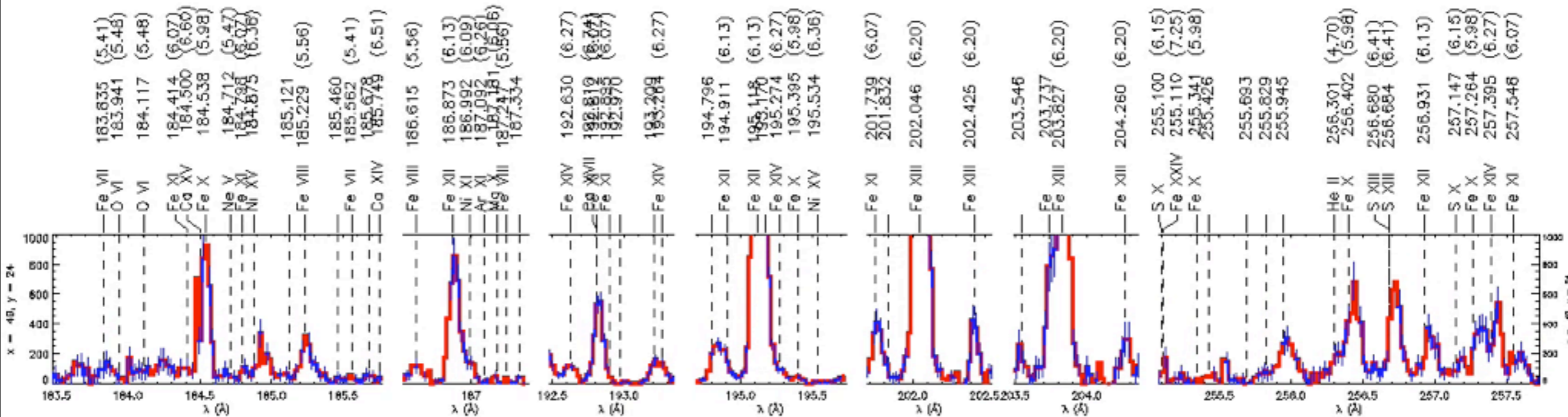
- 256.32 Å
- Only Doppler-shifted component is fitted here
- Clean area of spectrum
- In all cases, pattern persists for > 10 min where thread crosses slit

Multiple lines show blue-shift



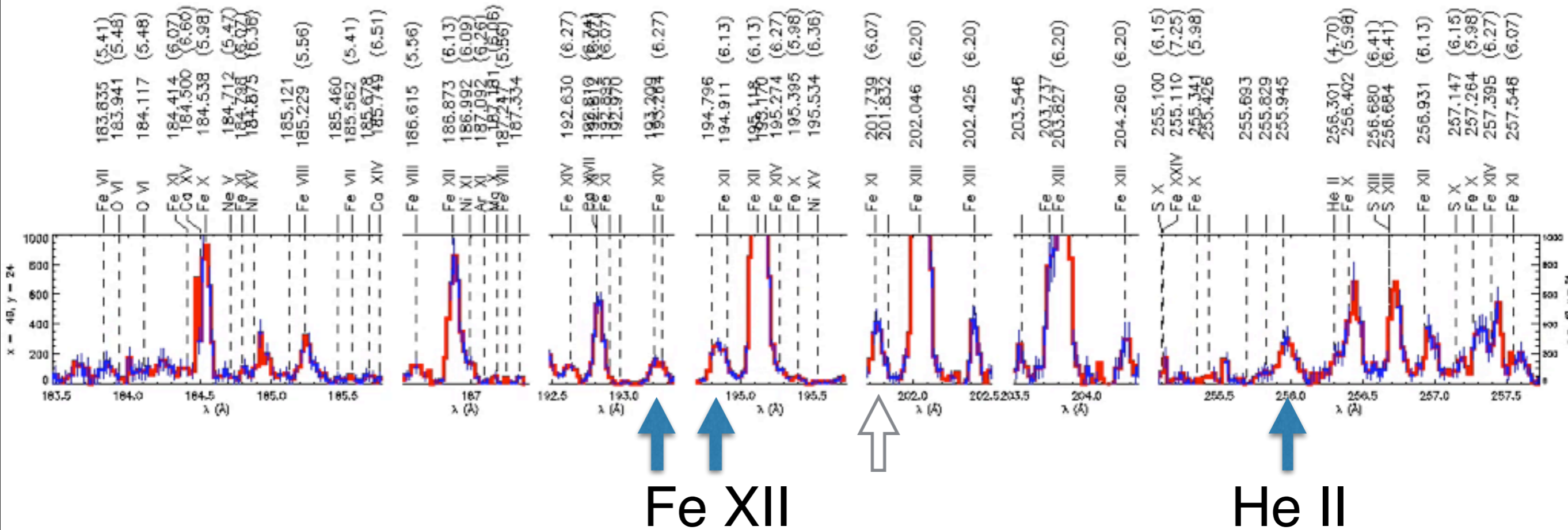
Selecting spectrum from increasingly higher points along slit as the movie plays

Multiple lines show blue-shift



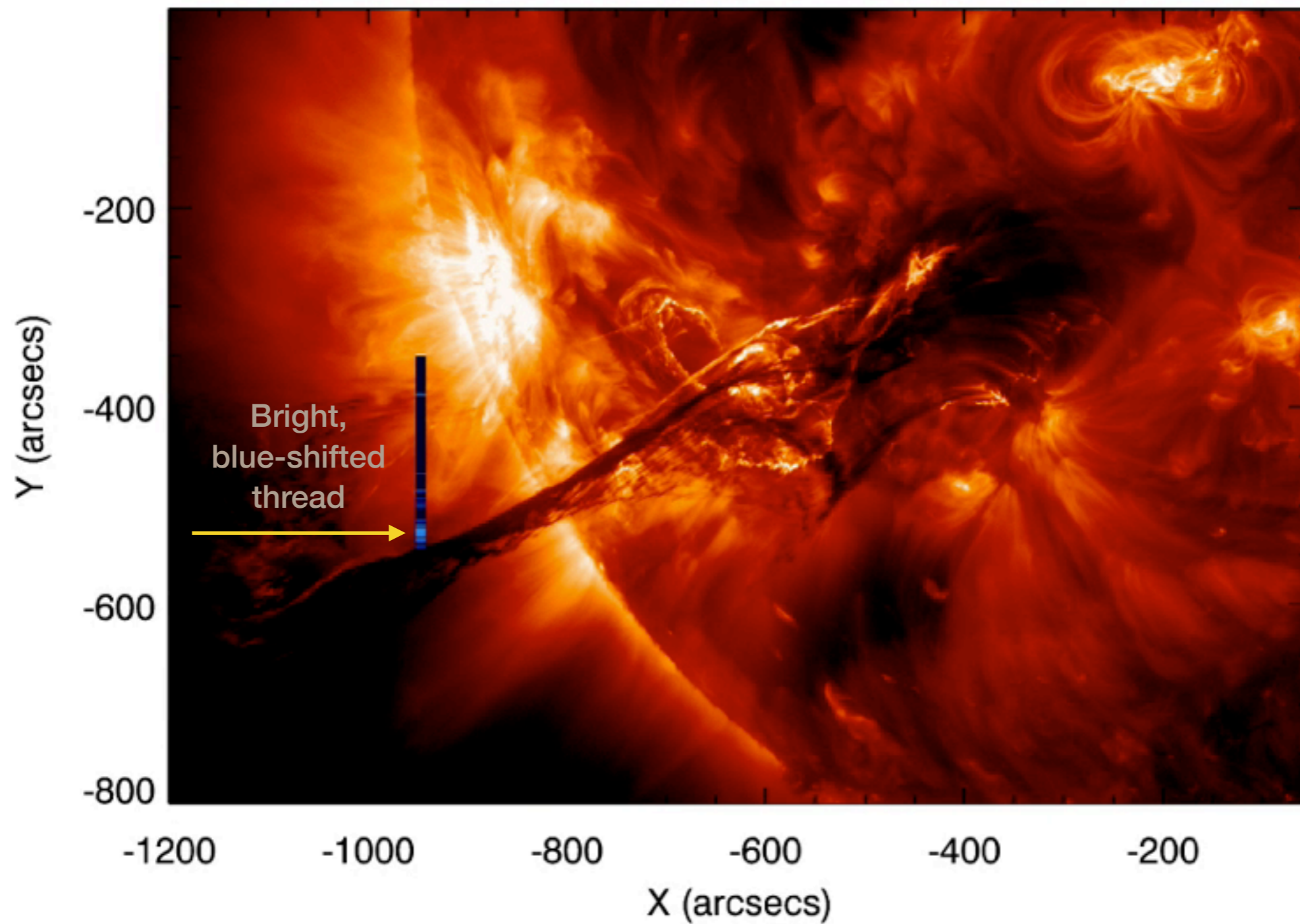
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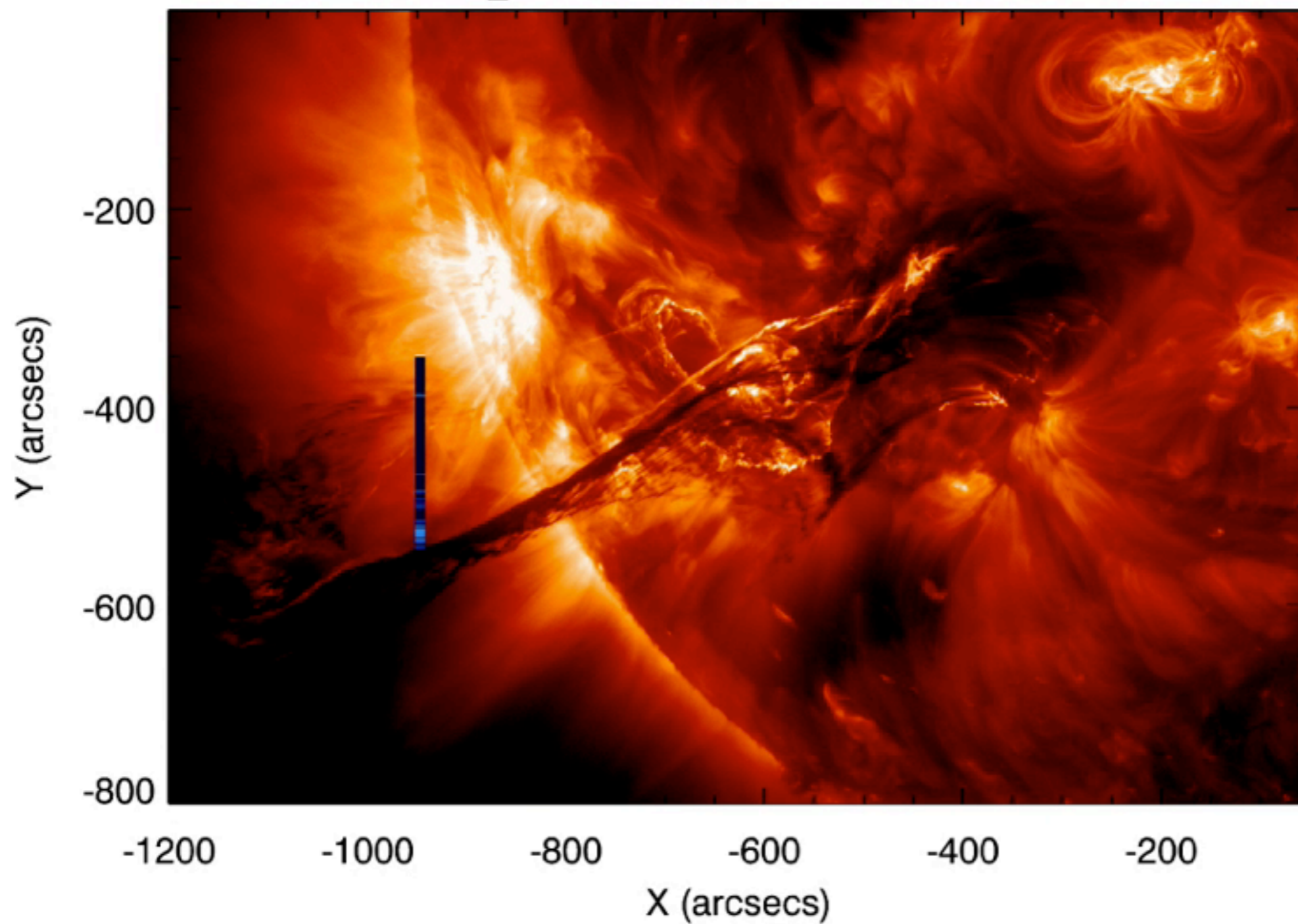


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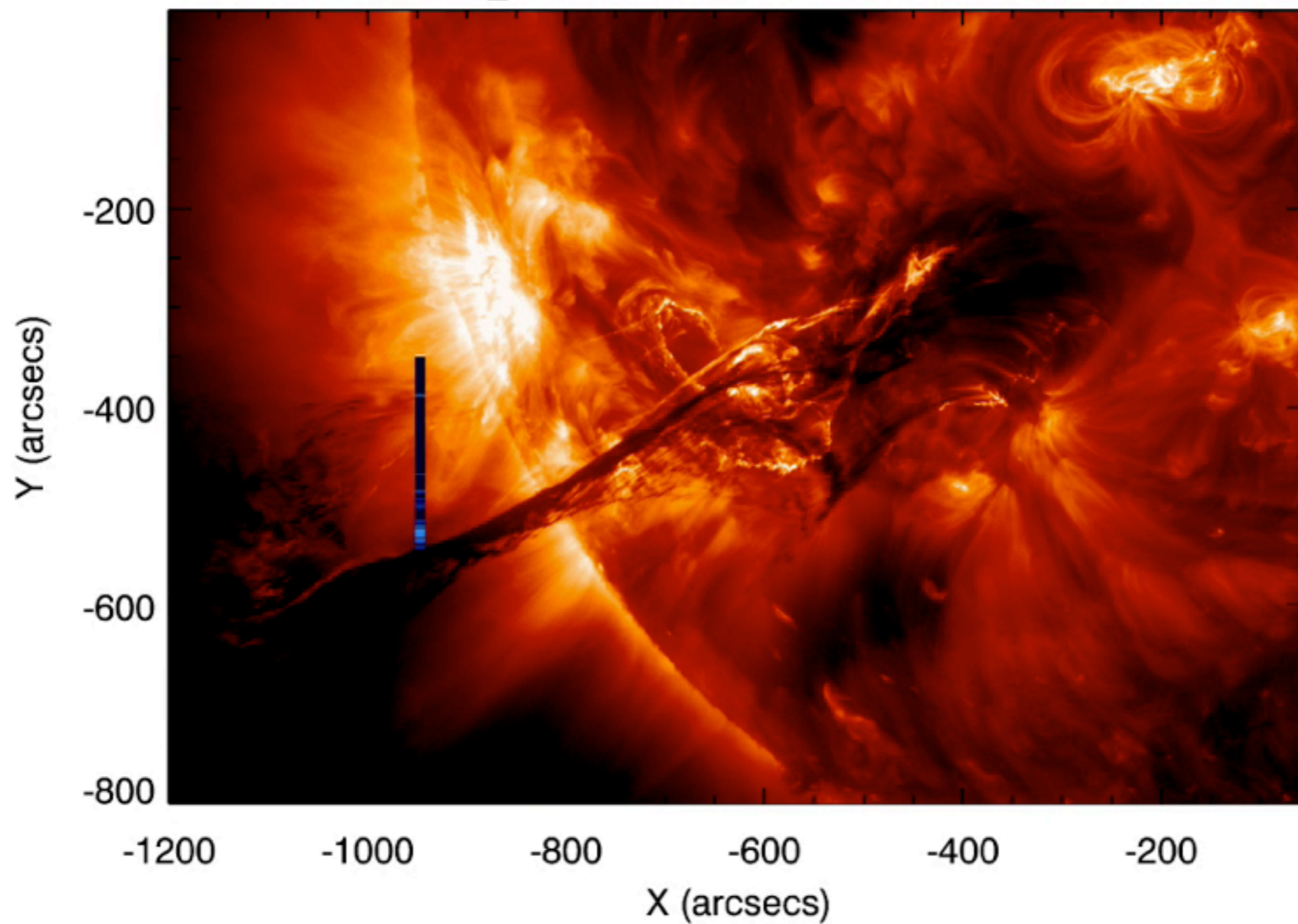
SDO AIA_2 193 31-Aug-2012 19:50:07.840 UT



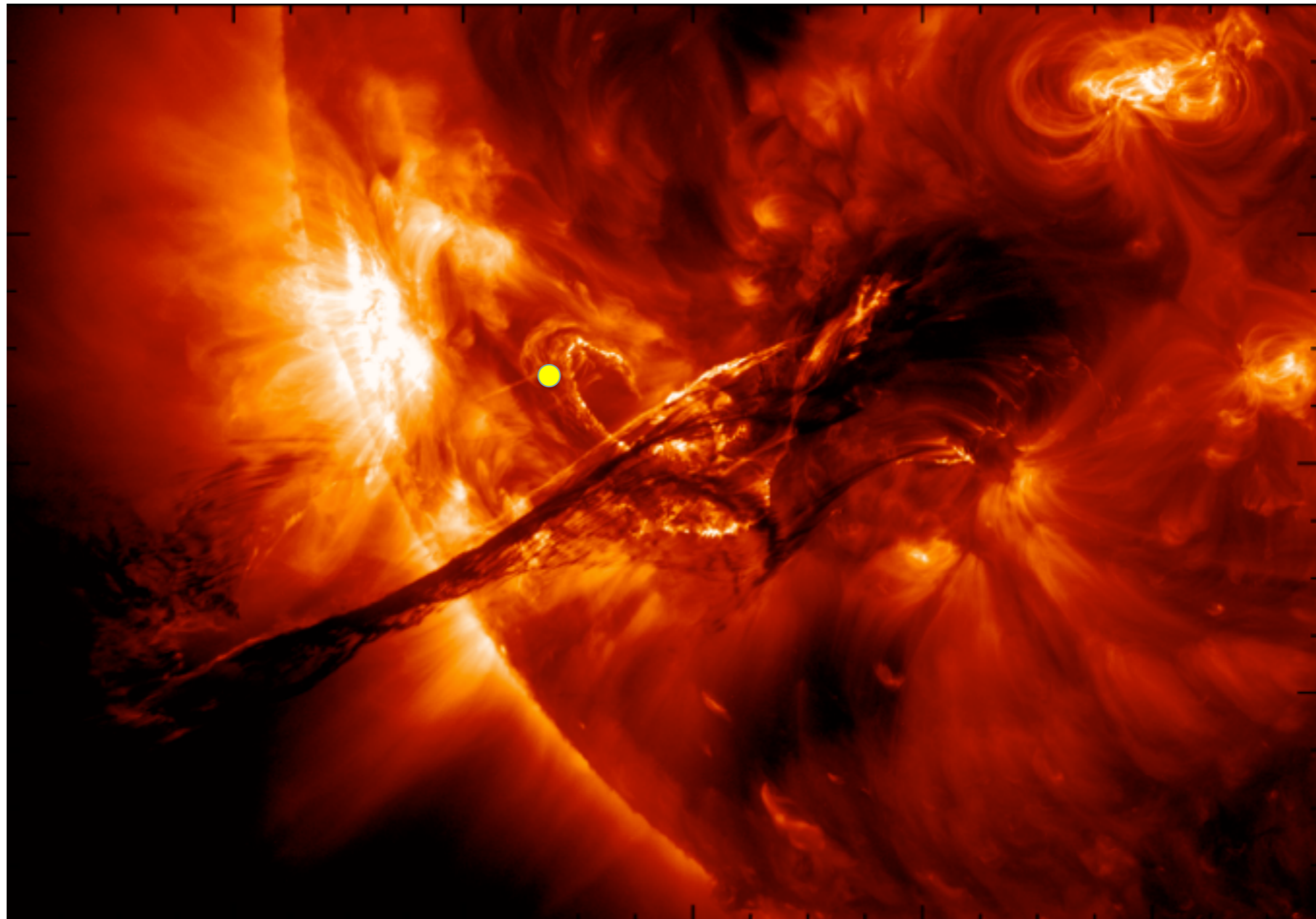
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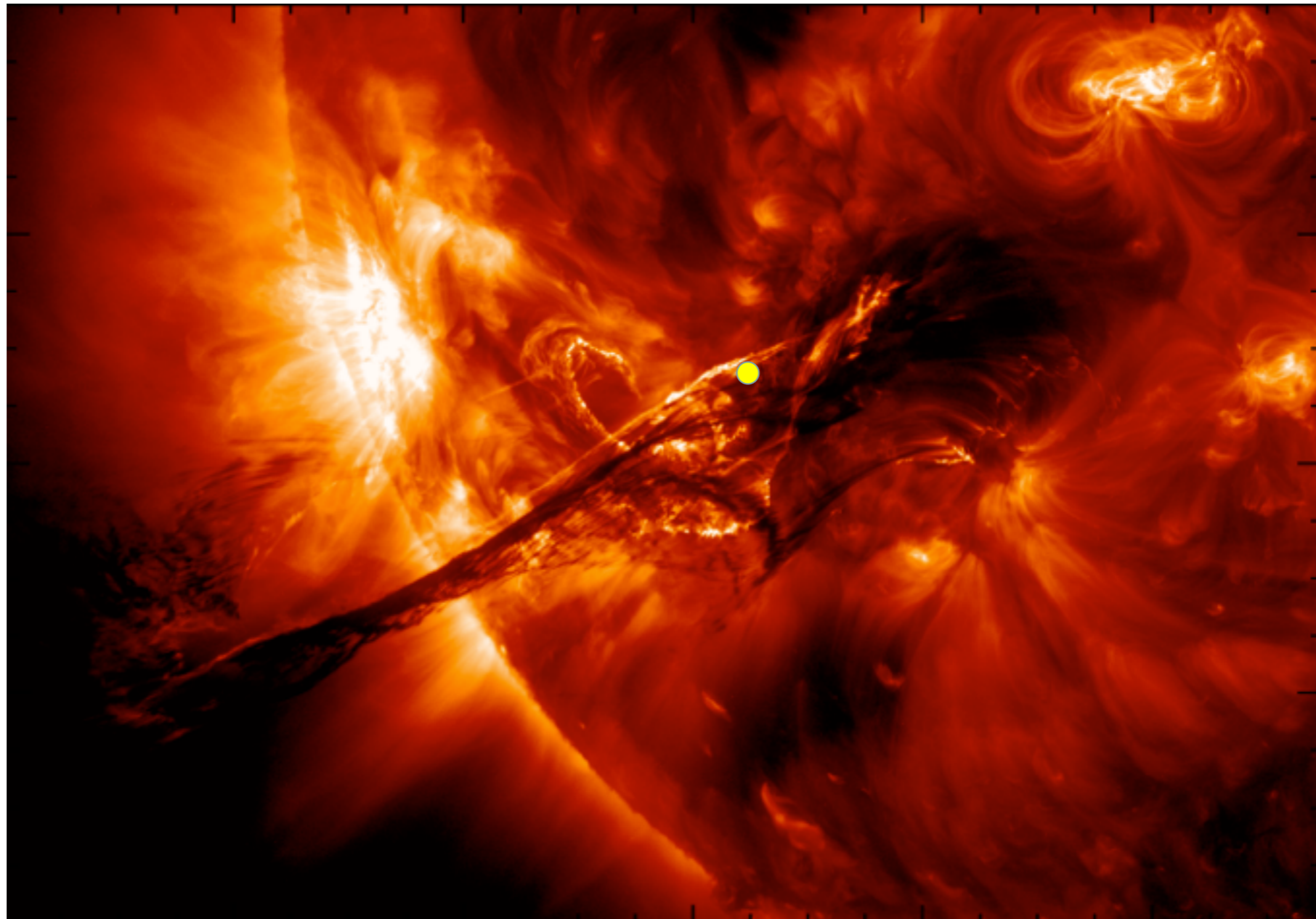
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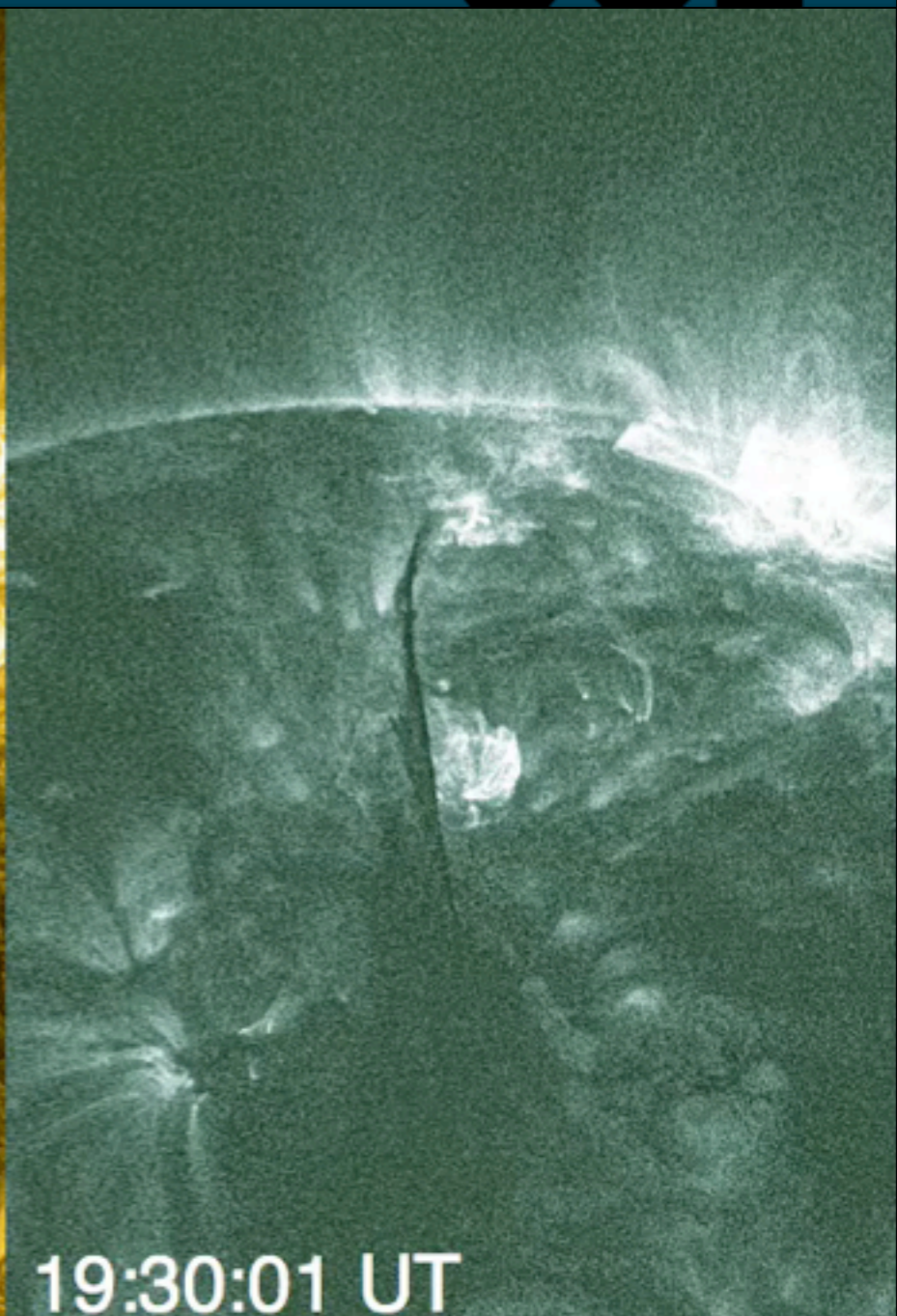
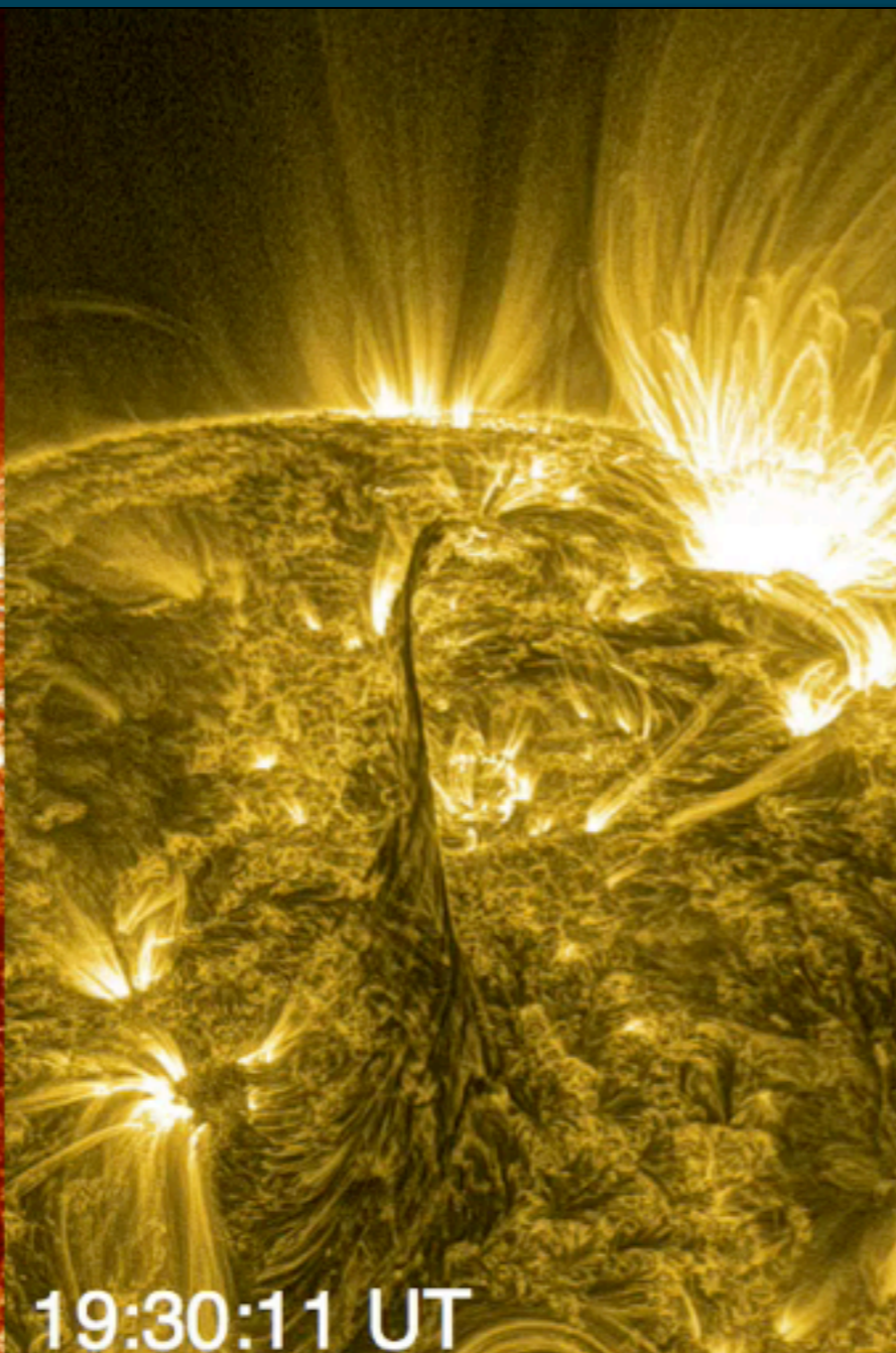


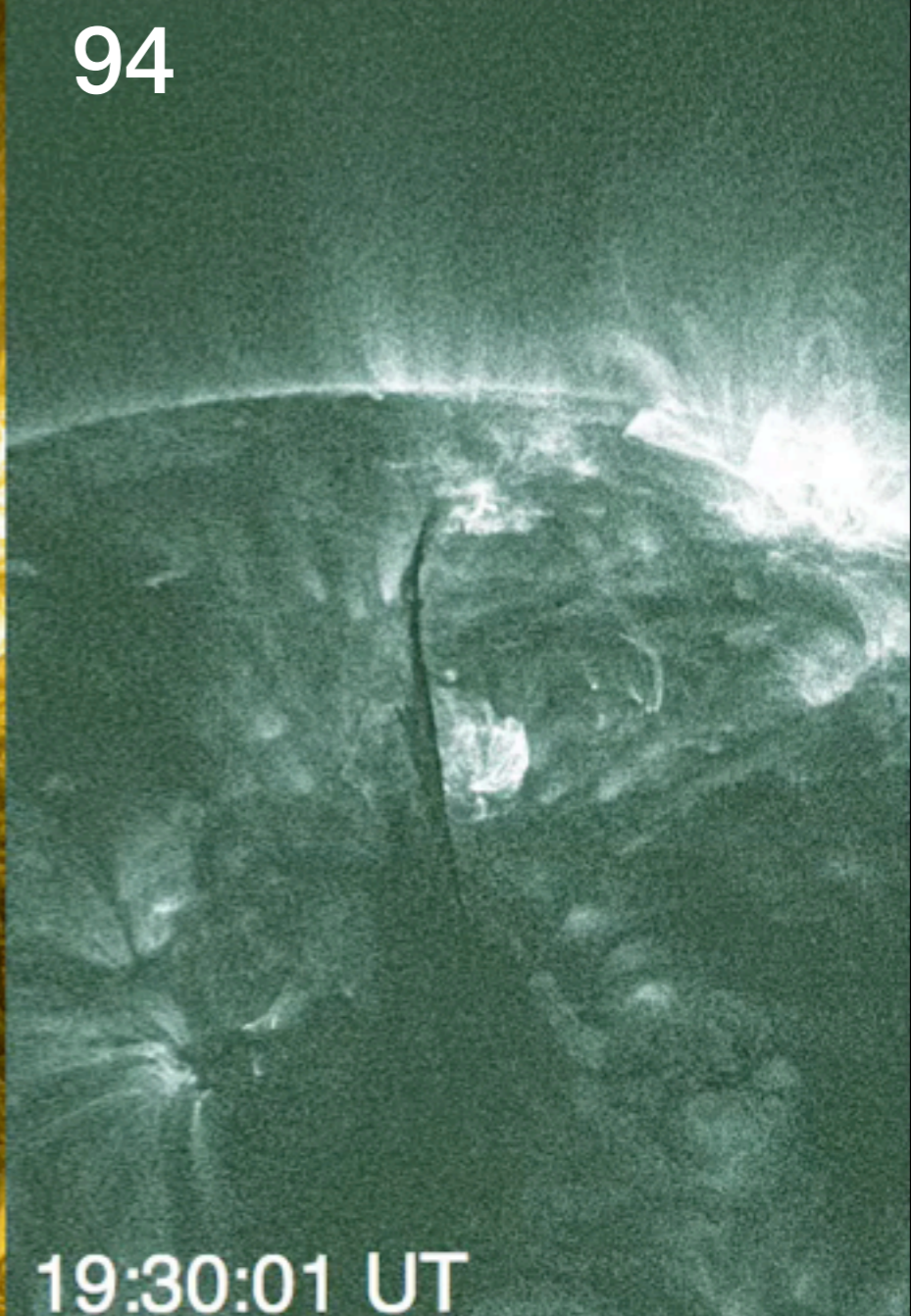
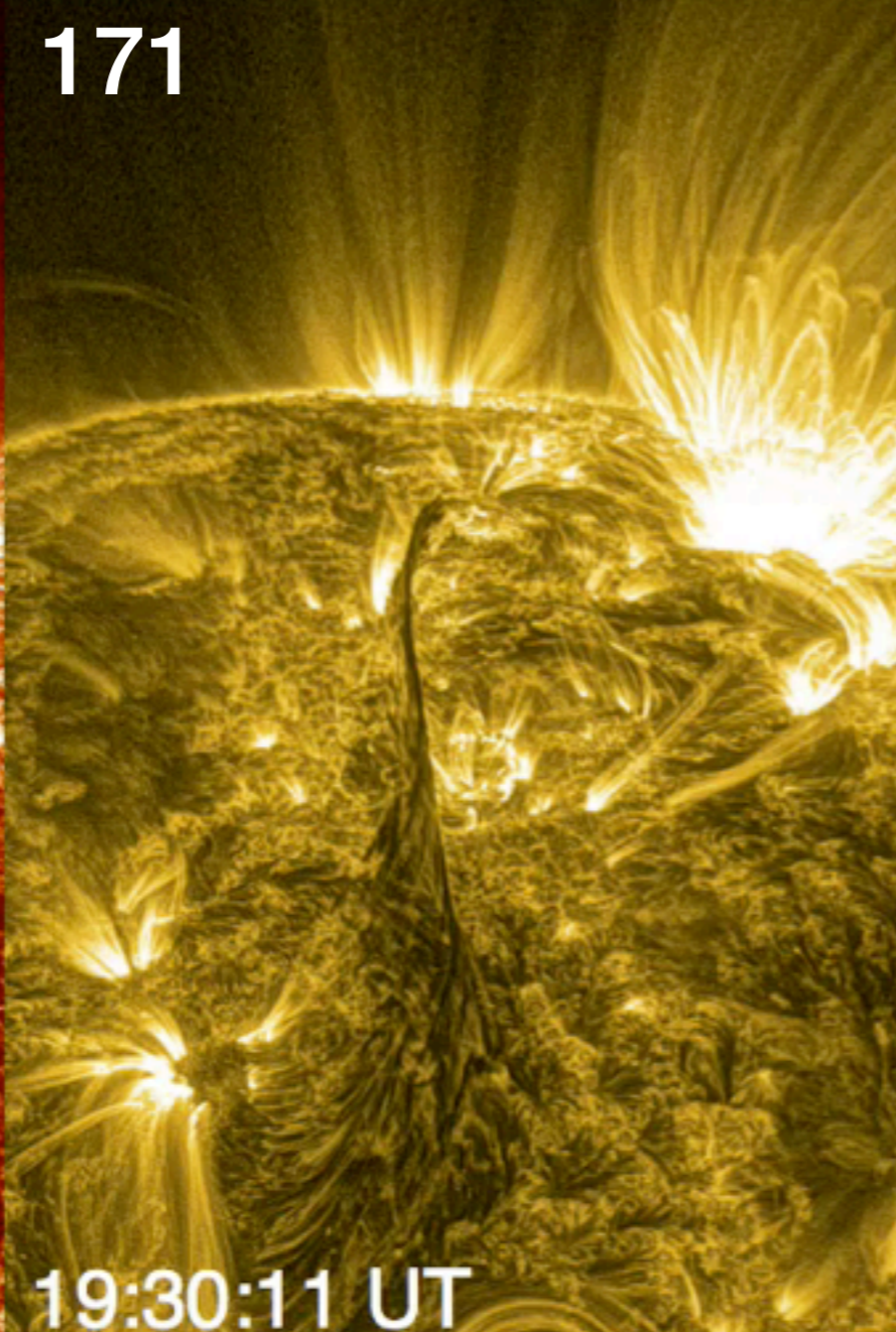
Thread outline



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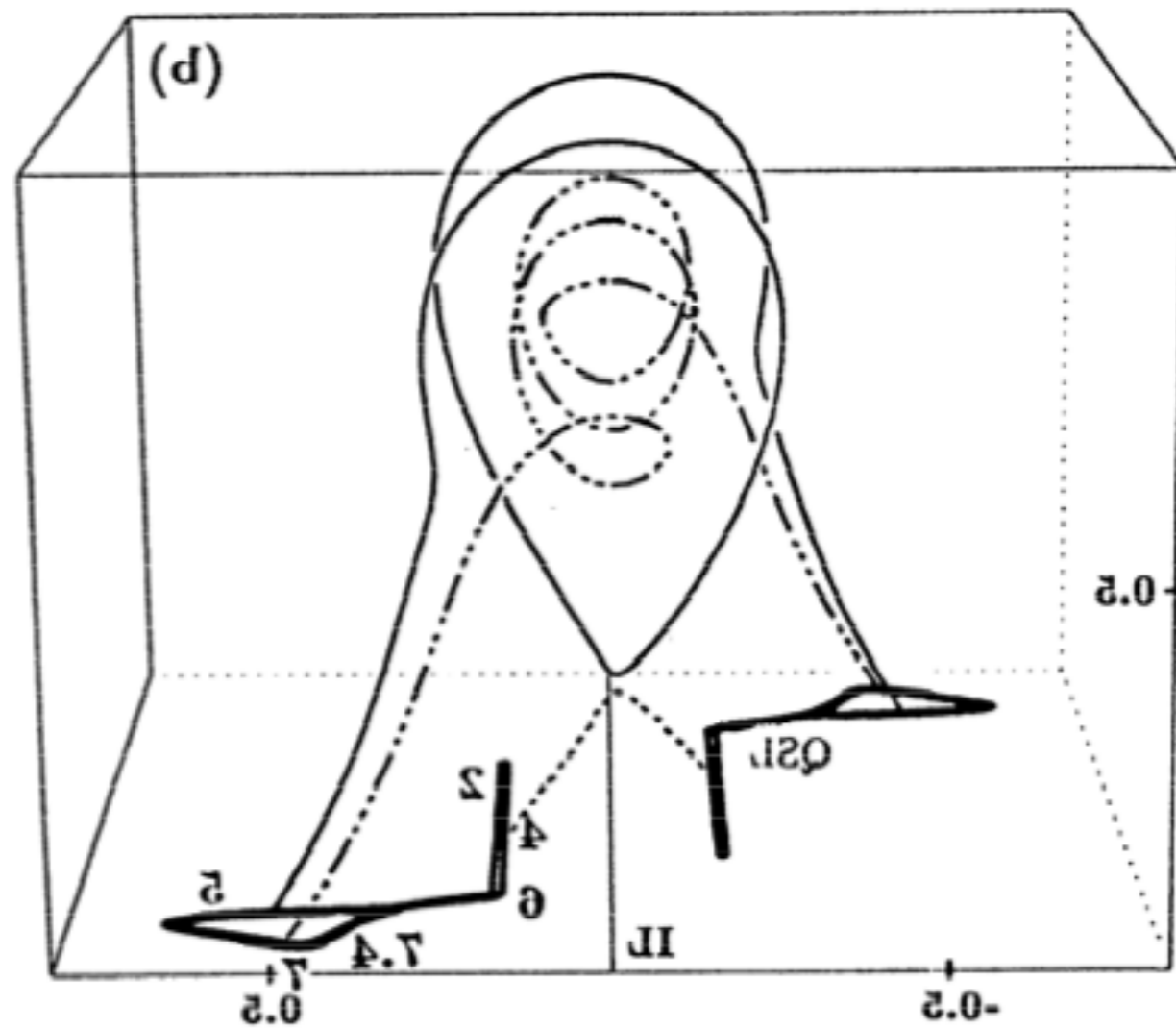




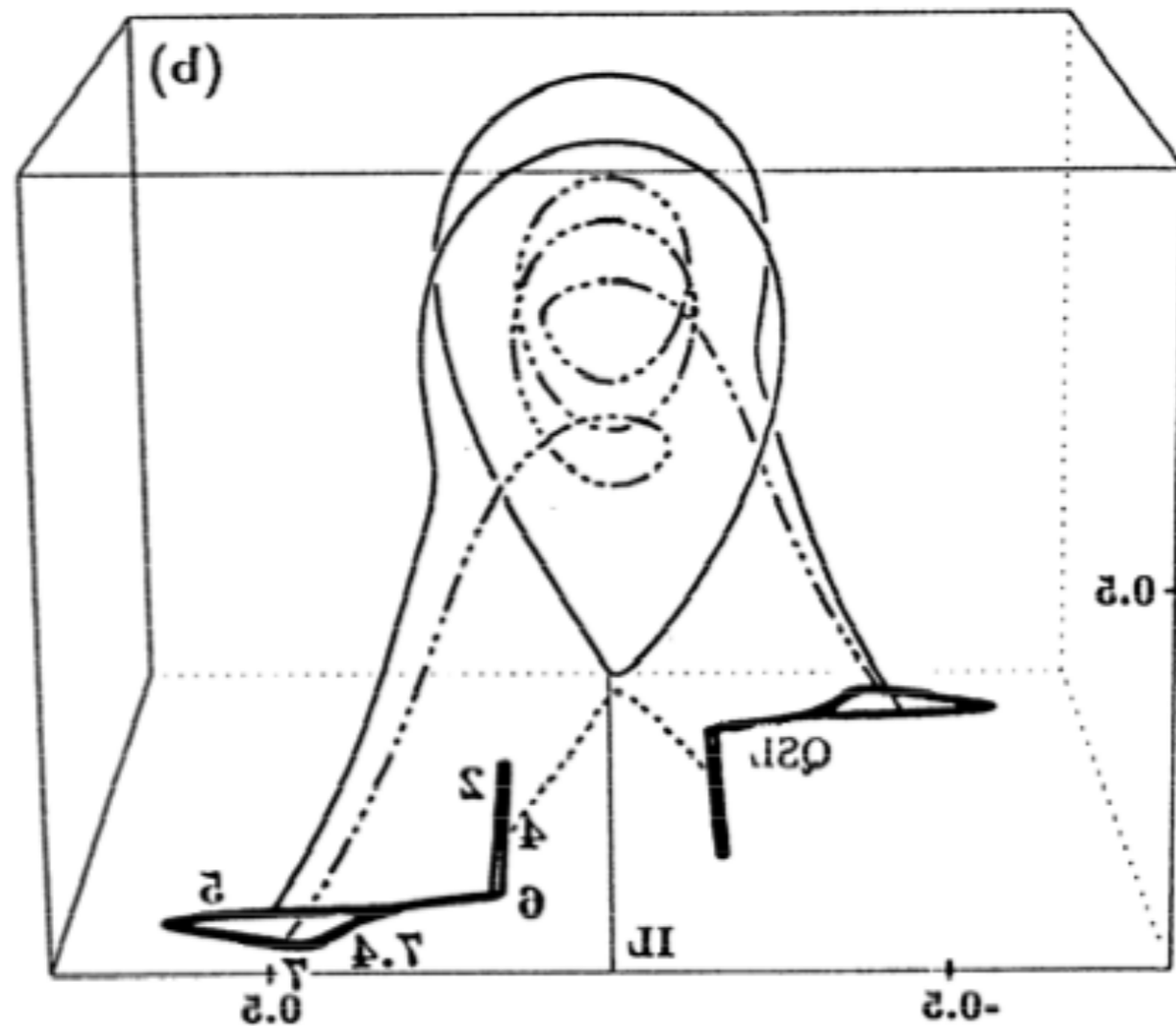
Flare ribbons & topology



Flare ribbons & topology

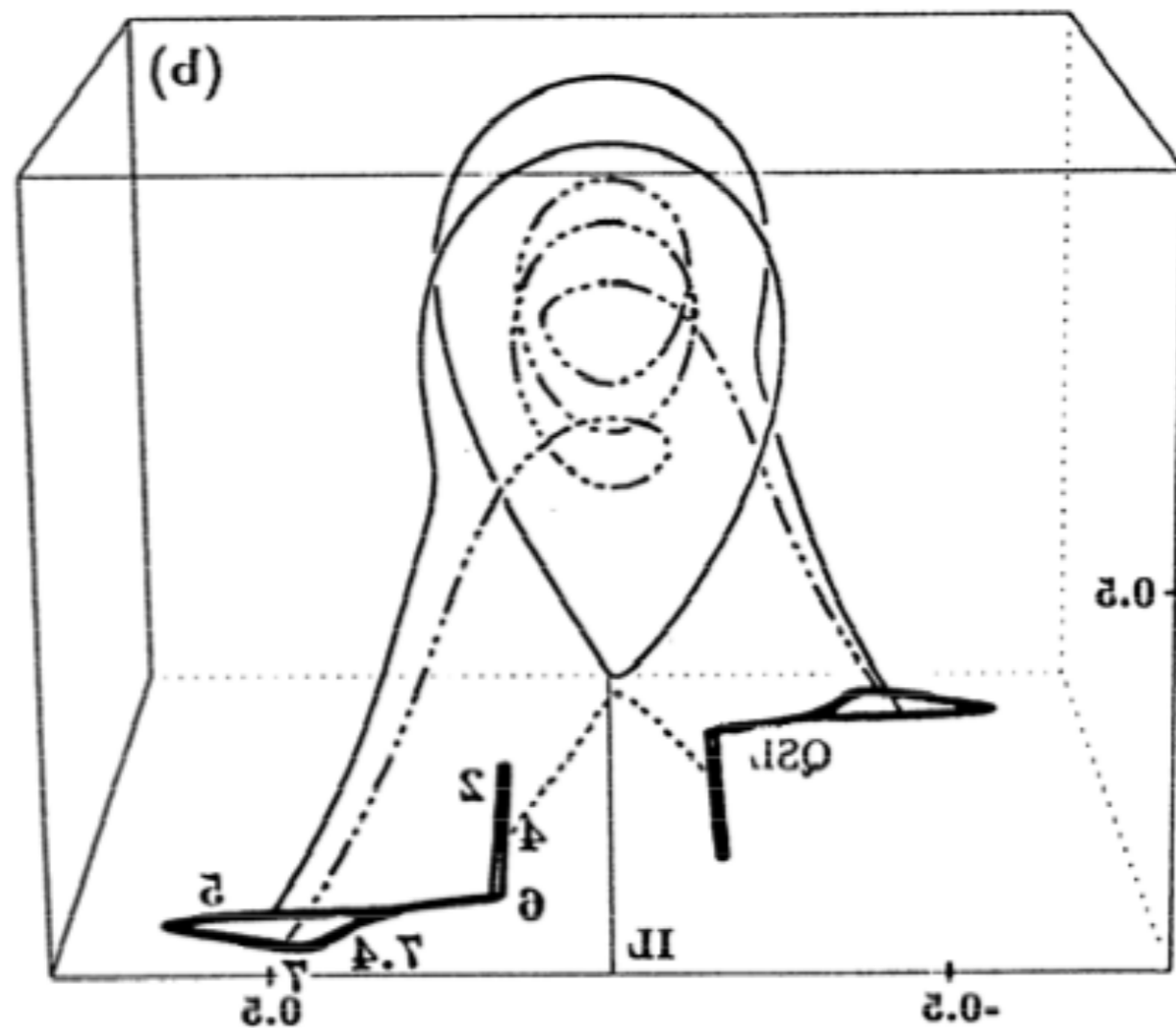


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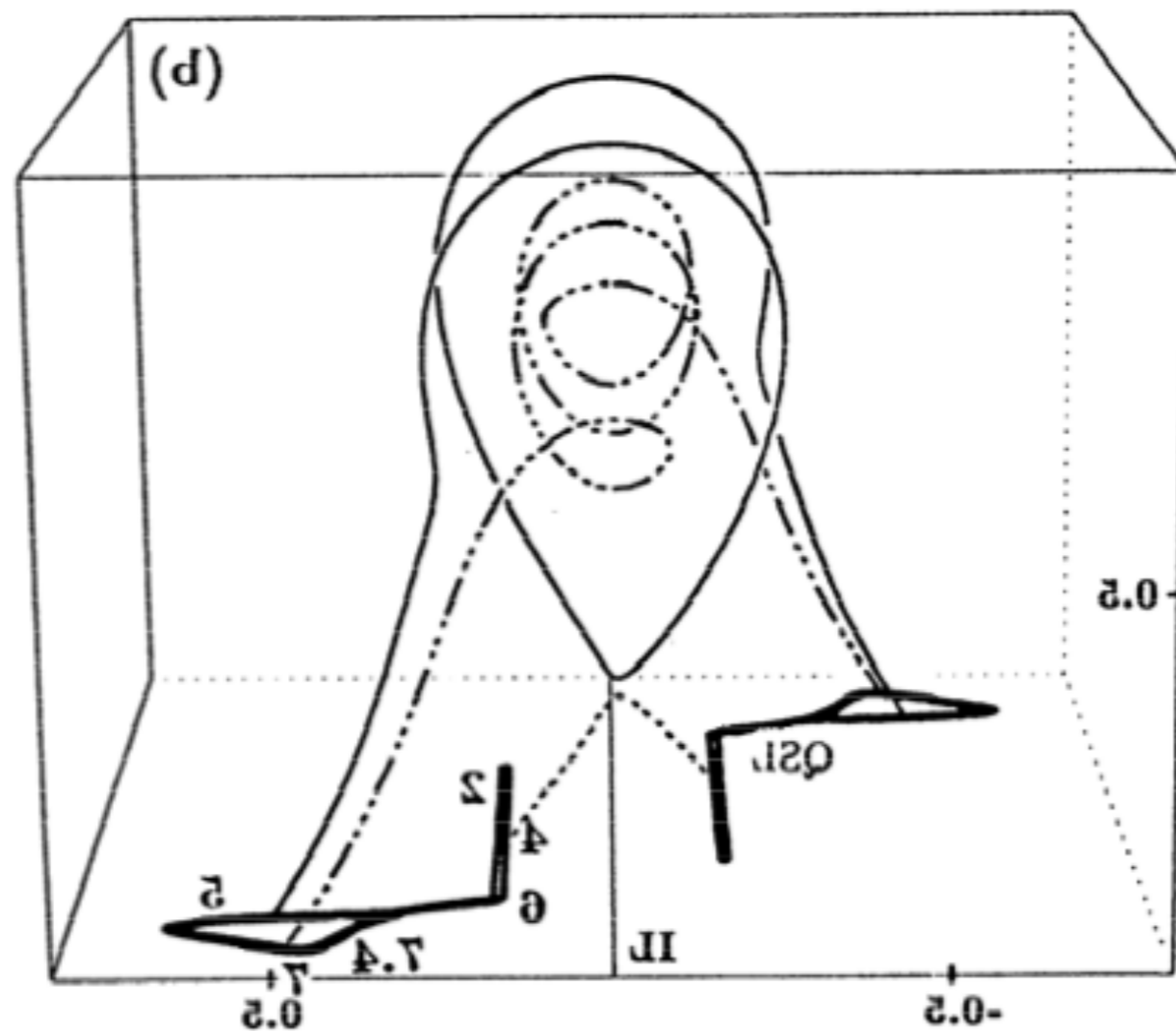
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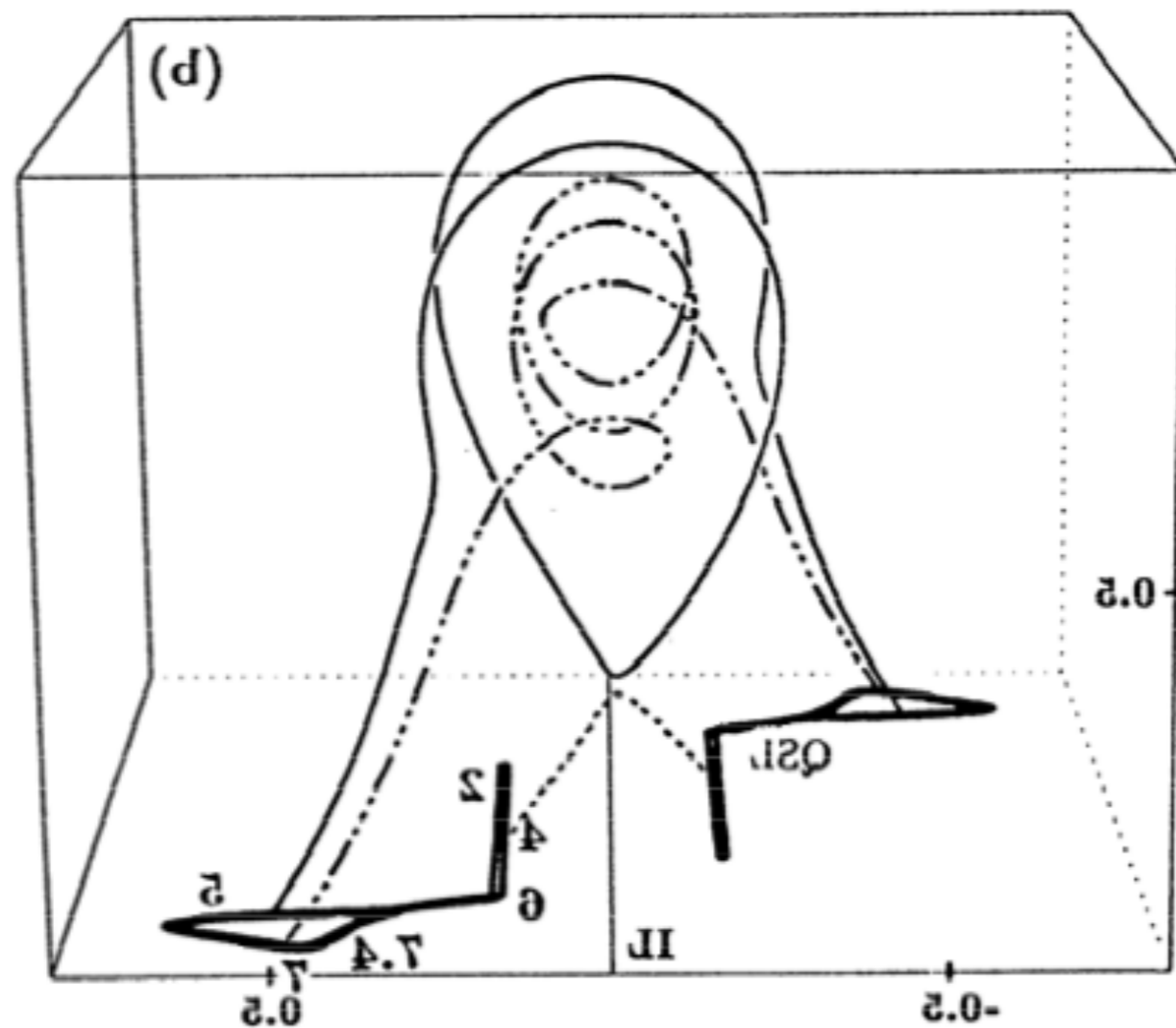
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- Flare ribbons represent intersection of QSL with photosphere

Flare ribbons & topology

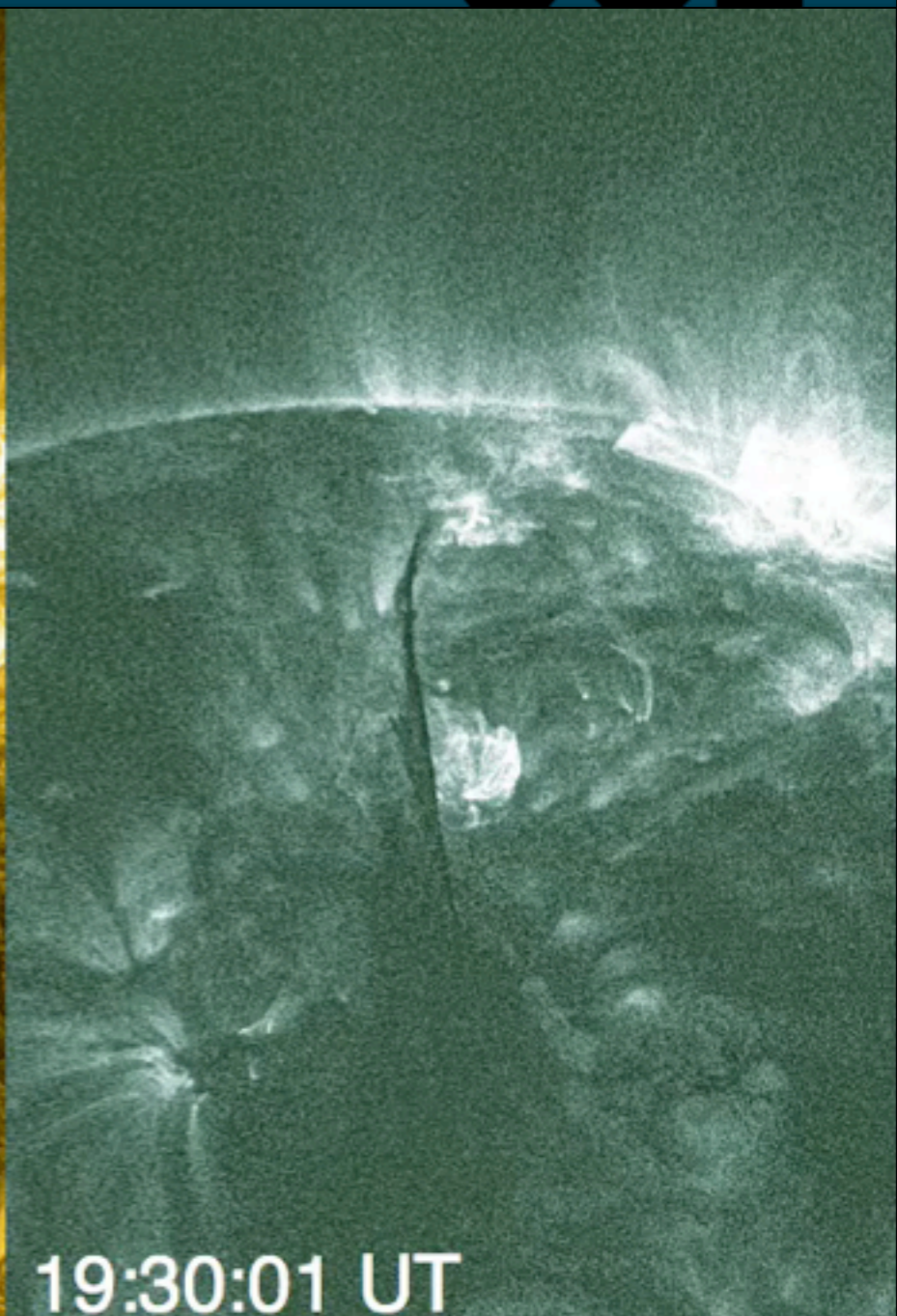
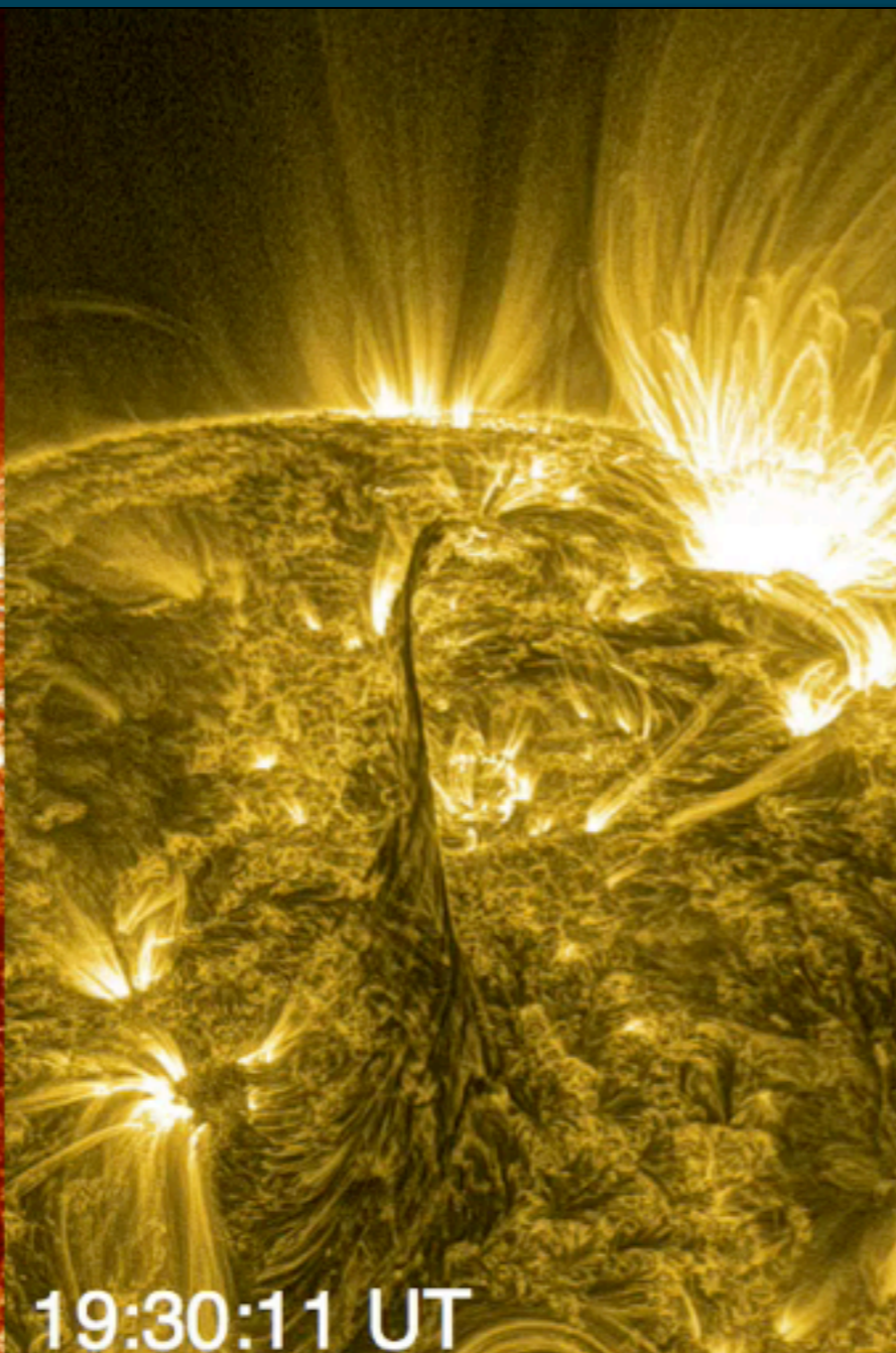


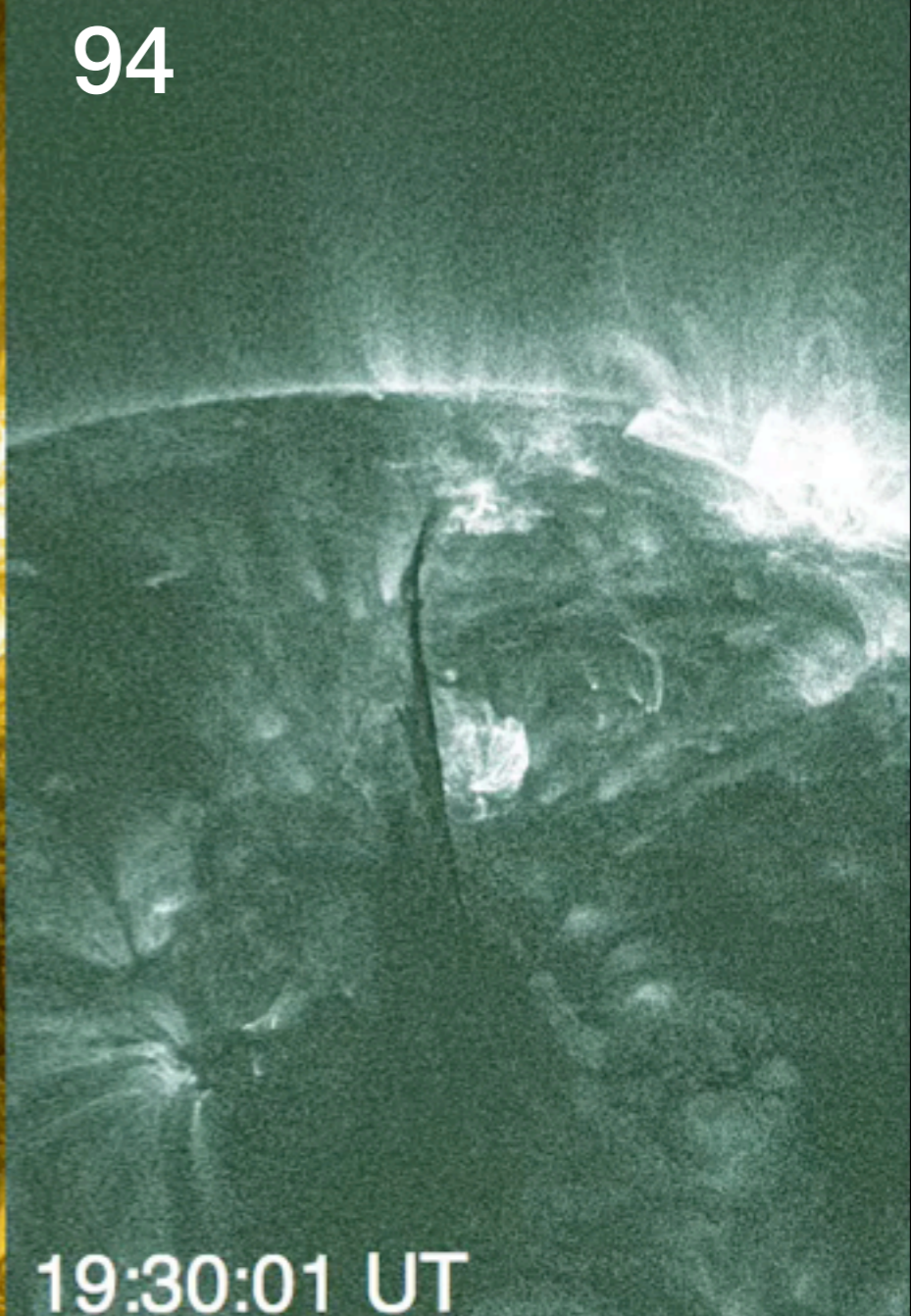
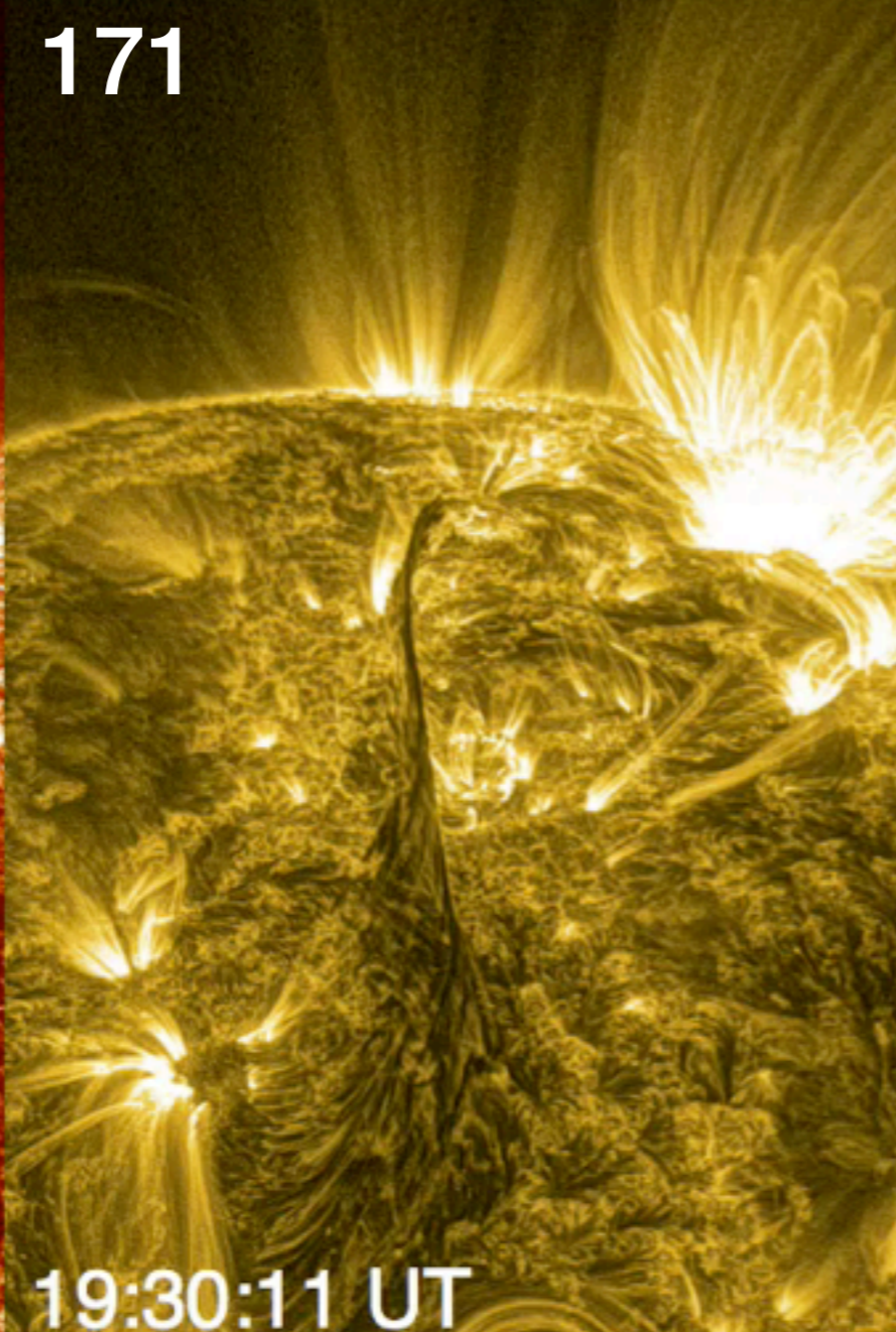
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Flare ribbons & topology



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- Flare ribbons represent intersection of QSL with photosphere
- Prime topological feature where reconnection can happen
- Anchoring of bright thread in/ surrounded by flare ribbons is hint that thread may be driven by *reconnection*





Thread flows & temperature

Thread flows & temperature

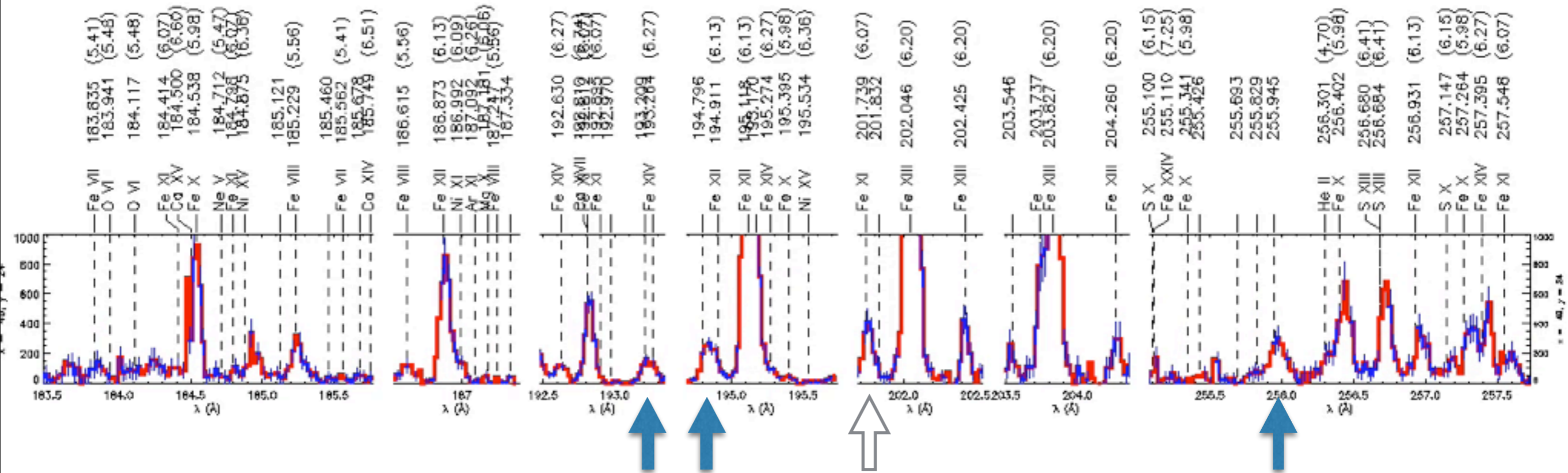
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Thread flows & temperature

- One thread lights up and contains rapidly accelerated plasma
- Does it heat up?
 - Doppler shifts indicate flows are supersonic unless $T \geq 8 \text{ MK}$.

But seen in “warm” lines only

Doppler motions in warm ($T \leq 2$ MK) lines, but no hotter.



Summary

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- The explanation for these up-flows has to fit into our picture of filament eruptions